# Foliations of small tubes in Riemannian manifolds by capillary minimal discs 

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July 17, 2008
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#### Abstract

Letting $\Gamma$ be an embedded curve in a Riemannian manifold $\mathcal{M}$, we prove the existence of minimal disc-type surfaces centered at $\Gamma$ inside surface of revolution of $\mathcal{M}$ around $\Gamma$, having small radius, and intersecting it with constant angles. In particular we obtain that small tubular neighborhoods can be foliated by minimal discs.


## 1 Introduction

Minimal surfaces are surfaces with mean curvature vanishing everywhere. These include, but are not limited to, surfaces of minimal area subjected to various constraints.
In this paper we are interested in minimal surfaces which intersect a given hyper-surface with a constant angle. Such surfaces, called capillary minimal surfaces, are critical points of an energy functional under some constraints, see (3) below.
Capillary surfaces correspond to the physical problem of describing of an incompressible liquid in a container in the absence of gravity. A great deal of work has been devoted to capillarity phenomena from the point of view of existence, uniqueness and topological properties of solutions mainly in the non-parametric case and in the more general situation of presence of gravity (see the book of R. Finn, [7], for an account of the subject). Some answers to these questions have been obtained by many authors. We refer for example to the papers [9], [10], [4] [12], [16], [18] [17], [20],,[21] and the references therein.
Here we prove existence results of capillary surfaces with prescribed topology in Riemannian manifolds. Roughly speaking, we first show the existence of a class of capilary (minimal) disc-type surfaces embedded in a Riemannian surface of revolution (see below). In particular, shrinking enough the thickness of the surface of revolution, this class constitutes a foliation. Secondly we have existence of minimal disc-type surfaces embedded in a geodesic tube of a curve which intersect perpendicularly the boundary of the tube.
The method we use is perturbative in nature and the main idea goes back to R. Ye, [23], subsequently employed with success by many authors to obtain existence of (large) constant mean curvature hyper-surfaces. For example, one can see [5], [6] [13], [14] and [15]. Before stating our results, some preliminaries are required.

A surface of revolution is a surface created by rotating a parametric curve $[a, b] \ni s \rightarrow$ $(\kappa(s), \phi(s)) \in \mathbb{R}^{2}$ lying on some plane around a straight line (the axis of rotation) in the same plane.
The resulting surface $\mathscr{C}^{1}$ therefore always has azimuthal symmetry. Examples of surfaces of revolution include cylinder (excluding the ends), hyperboloid, paraboloid, sphere, torus, etc.
In more generality one can obtain surfaces of revolution in $\mathbb{R}^{n+1}, n \geq 2$ using the standard parametrization

[^0]$$
S(s, z)=(\kappa(s), \phi(s) \Theta(z)),
$$
where $z \mapsto \Theta(z) \in S^{n-1}, \phi(s) \neq 0 \quad \forall s \in[a, b]$.
Assuming that the rotating curve is parameterized by arc length namely
$$
\left(\phi^{\prime}(s)\right)^{2}+\left(\kappa^{\prime}(s)\right)^{2}=1
$$
clearly the disc $\mathscr{D}_{s, 1}$ centered at $(\kappa(s), 0)$ (on the axis of rotation) with radius $\phi(s)$ parameterized by
$$
B_{1}^{n} \ni x \mapsto(\kappa(s), \phi(s) x),
$$
has zero mean curvature and intersects the above surface of revolution with a constant angle equal to $\arccos \phi^{\prime}(s)$, where $B_{1}^{n}$ stands for the unit ball of $\mathbb{R}^{n}$ centered at the origin, namely $\mathscr{D}_{s, 1}$ is a capillary surface.
Motivated by capillarity problems, for questions of stability, see [7], it is not restrictive to assume that the angle of contact is in $(0, \pi)$, namely $\phi^{\prime}(s) \in(-1,1)$ or equivalently
$$
\kappa^{\prime}(s) \neq 0
$$

We shall extend these definitions of surface of revolution in a Riemannian setting.
Let $\left(\mathcal{M}^{n+1}, g\right)$ be Riemannian manifold, and $\Gamma$ an embedded curve parameterized by a map $\gamma:[0,1] \rightarrow \mathcal{M}$. We consider a local parallel orthogonal frame $E_{1}, \cdots, E_{n}$ of $N \Gamma$ along $\Gamma$. This determines a coordinate system by

$$
[0,1] \times \mathbb{R}^{n} \ni\left(x_{0}, y\right) \mapsto f\left(x_{0}, y\right):=\exp _{\gamma\left(x_{0}\right)}\left(y^{i} E_{i}\right) \in \mathcal{M}
$$

For a small parameter $\rho>0$, consider the Riemannian surface of revolution $\mathscr{C}^{\rho}$ around $\Gamma$ in $\mathcal{M}$ parameterized by

$$
(s, z) \longrightarrow f(\rho S(s, z))=f(\rho \kappa(s), \rho \phi(s) \Theta(z))=\exp _{\gamma(\rho \kappa(s))}\left(\rho \phi(s) \Theta^{i}(z) E_{i}\right)
$$

where $z \mapsto \Theta(z) \in S^{n-1}$, and call its interior $\Omega_{\rho}:=\operatorname{int} \mathscr{C}^{\rho}$ which is nothing but a tubular neighborhood for $\Gamma$ if $\rho$ is small enough. Here we are assuming always that $\phi(s) \neq 0$ and that $\left(\phi^{\prime}(s)\right)^{2}+\left(\kappa^{\prime}(s)\right)^{2}=1$.

For any $s \in[a, b]$, we consider the following set

$$
D_{s, \rho}:=f\left(\rho \kappa(s), \rho \phi(s) B_{1}^{n}\right)
$$

it is clear that $\partial D_{s, \rho} \subset \mathscr{C}^{\rho}$ and we have that the mean curvature $H_{D_{s, \rho}}$ of $D_{s, \rho}$, see $\S 4$ 4.1, satisfies

$$
\begin{equation*}
H_{D_{s, \rho}}=\mathcal{O}(\rho) \quad \text { in } D_{s, \rho} \tag{1}
\end{equation*}
$$

while the angle between the unit outer normals (see also § 4.2) can be expanded as

$$
\begin{equation*}
\left\langle N_{D_{s, \rho}}, N_{\mathscr{C} \rho}\right\rangle=\phi^{\prime}(s)+\mathcal{O}(\rho) \quad \text { on } \partial D_{s, \rho} \tag{2}
\end{equation*}
$$

Our aim is to perturb $D_{s, \rho}$ to a capillary minimal submanifold, $\mathscr{D}_{s, \rho}$, of $\Omega_{\rho}$ centered on $\Gamma$ with contact angle $\arccos \phi^{\prime}(s)$ along $\partial \mathscr{D}_{s, \rho} \subset \mathscr{C}^{\rho}$, as it happens in $\mathbb{R}^{n+1}$.
Theorem 1.1 Suppose we are in the situation described above. Let $\left[a^{\prime}, b^{\prime}\right] \subset[a, b]$ be such that $\phi(s) \phi^{\prime \prime}(s)>0$ for every $s \in\left[a^{\prime}, b^{\prime}\right]$. Then there exists $\rho_{0}>0$ such that for any $s \in\left[a^{\prime}, b^{\prime}\right]$ and $\rho \in\left(0, \rho_{0}\right)$, there exists an embedded minimal disc $\mathscr{D}_{s, \rho} \subset \Omega_{\rho}$, intersecting $\mathscr{C}^{\rho}$ by an angle equal to $\phi^{\prime}(s)$ along its boundary. Moreover $\mathscr{D}_{s, \rho}$ is a normal graph over the set $D_{s, \rho}$ for which the norm (in the $\mathcal{C}^{2, \alpha}$-topology) of this function defining the graph tends to zero uniformly as $\rho$ tends to zero.
Furthermore there exists a tubular neighborhood $O_{\rho}$ of $\gamma\left(\left[a^{\prime}, b^{\prime}\right]\right)$ foliated by such minimal discs for which each leaf intersects $\partial O_{\rho}$ transversally along its boundary.

Remark 1.2 - When we parameterize in particular $\mathscr{C}^{\rho}$ with $\kappa(s)=s$, and if we require the capillary discs to be perpendicular to $\mathscr{C}^{\rho}$, we obtain the conditions $\phi^{\prime}=0$ and $\phi^{\prime \prime} \neq 0$. This means that non-degenerate extrema of the width $\phi$ determine the location of such surfaces.

- An example is the hyperboloid, $\phi(s)=\cosh s$ and $\kappa(s)=\sinh s$. Here one may see $\mathcal{M}$ as a Lorentzian manifold modeled on the Minkowski space $\mathbb{R}_{1}^{n}$. Letting $q \in \mathcal{M}$ and $E_{0}$ a unit time-like vector of $T_{q} \mathcal{M}$ and $\gamma\left(x_{0}\right)=\exp _{q}\left(x_{0} E_{0}\right)$ so one can see $\mathscr{D}_{0, \rho}$ as a space-like minimal disc in the geodesic sphere of radius $\rho$.

An interesting particular case which is not covered by Theorem 1.1 is when $\phi \equiv 1$ and $\kappa=\operatorname{Id}$, namely when we deal with geodesic tubes. In this situation (recall that in this case the angle of contact is $\frac{\pi}{2}$ ) it is the geometry of the manifold to determine the position of the discs. More precisely, we have that $\mathscr{C}^{\rho}$ is the geodesic tube of radius $\rho>0$ around $\Gamma$,

$$
\mathscr{C}^{\rho}=\left\{q \in \mathcal{M} \quad: \quad \operatorname{dist}_{g}(q, \Gamma)=\rho\right\},
$$

and its interior is nothing but

$$
\Omega_{\rho}:=\left\{q \in \mathcal{M} \quad: \quad \operatorname{dist}_{g}(q, \Gamma)<\rho\right\} .
$$

In this case due to invariance by translations along the axis of rotation, we reduced our problem of finding minimal surfaces to a finite-dimensional one. Namely we have obtained the following

Theorem 1.3 There exists a smooth function $\psi_{\rho}:[a, b] \rightarrow \mathbb{R}$ such that, for $\rho$ small, if $s_{0}$ is a critical point of $\psi_{\rho}$ the set $D_{s_{0}, \rho}$ can be smoothly perturbed to an embedded minimal hyper-surface $\mathscr{D}_{s_{0}, \rho} \subset \Omega_{\rho}$ intersecting $\mathscr{C}^{\rho}$ perpendicularly along its boundary. Furthermore, for any integer $k$, there exists a constant $c_{k}$ (independant on $\rho$ ) such that

$$
\left\|\psi_{\rho}-\sum_{i, j}^{n}\left\langle R_{p}\left(E_{j}, E_{i}\right) E_{j}, E_{i}\right\rangle\right\|_{\mathcal{C}^{k}[a, b]} \leq c_{k} \rho^{2}
$$

where $R_{p}$ is the Riemann tensor of $\mathcal{M}$ at $p=\gamma(\rho s)$.
Some remarks are due: let $\Gamma \ni p \rightarrow \Psi(p)=\sum_{i, j}^{n}\left\langle R_{p}\left(E_{j}, E_{i}\right) E_{j}, E_{i}\right\rangle$ any strict maxima or minima of $\Psi$ imply the existence of minimal surfaces. In particular suppose at some point $p_{0}=\gamma\left(\rho s_{0}\right)$ interior to $\Gamma$, there hold

$$
d \Psi\left(p_{0}\right)\left[\dot{\gamma}\left(\rho s_{0}\right)\right]=0 \quad \text { and } \quad\left|d^{2} \Psi\left(p_{0}\right)\left[\dot{\gamma}\left(\rho s_{0}\right), \dot{\gamma}\left(\rho s_{0}\right)\right]\right|>c,
$$

for some constant $c$ independent on $\rho$. By the implicit function theorem, there exits a curve $\left(0, \rho_{0}\right) \ni \rho \mapsto s_{\rho}$ with $s_{\rho} \rightarrow s_{0}$ such that $s_{\rho}$ is a critical point of $\psi_{\rho}$. Hence for every $\rho \in\left(0, \rho_{0}\right)$, there exits an embedded minimal disc $\mathscr{D}_{s_{\rho}, \rho}$, centered at $\gamma\left(\rho s_{\rho}\right)$, contained in $\Omega_{\rho}$ that intersects $\partial \Omega_{\rho}$ perpendicularly along its boundary.

Remark 1.4 - We have that

$$
\Psi(p)=\sum_{i, j}^{n}\left\langle R_{p}\left(E_{j}, E_{i}\right) E_{j}, E_{i}\right\rangle=\boldsymbol{S}(p)+2 \operatorname{Ric}_{p}(\dot{\gamma}(\rho s), \dot{\gamma}(\rho s)),
$$

where

$$
\boldsymbol{S}(p)=\sum_{\alpha, \beta=0}^{n}\left\langle R_{p}\left(E_{\alpha}, E_{\beta}\right) E_{\alpha}, E_{\beta}\right\rangle
$$

is the scalar curvature of $\mathcal{M}$ at $p=\gamma(\rho s), E_{0}=\dot{\gamma}(\rho s)$ and Ric is the Ricci tensor of $\mathcal{M}$ at p. From Theorem 1.3, we have that if $s \mapsto \operatorname{Ric}_{p}(\dot{\gamma}(\rho s), \dot{\gamma}(\rho s))$ is locally constant along $\Gamma$ then stable critical points the scalar curvature yields existence of minimal discs.

- Recall that if $\left(\mathcal{M}_{1}^{m_{1}}, g_{1}\right)$ and $\left(\mathcal{M}_{2}^{m_{2}}, g_{2}\right)$ are two manifolds, the Riemann tensor $R$ of the (Riemannian) Cartesian product $\mathcal{M}^{m_{1}+m_{2}}:=\left(\mathcal{M}_{1} \times \mathcal{M}_{2}, g_{1} \oplus g_{2}\right)$ decomposes as $R=$ $R^{1} \oplus R^{2}$ since the connection $\nabla$ is given by $\nabla_{X_{1}+X_{2}}\left(Y_{1}+Y_{2}\right)=\nabla_{X_{1}}^{1} Y_{1}+\nabla_{X_{2}}^{2} Y_{2}$ for any $X_{1}, Y_{1}$ (resp. $X_{2}, Y_{2}$ ) vector fields of $\mathcal{M}_{1}$ (resp. $\mathcal{M}_{2}$ ), where $\nabla^{i}$ is the connection of $\mathcal{M}_{i}$. Clearly for any $p_{2} \in \mathcal{M}_{2}$, the set $\left(\mathcal{M}_{1}\right)\left(p_{2}\right):=\left\{\left(p_{1}, p_{2}\right) \in \mathcal{M}: p_{1} \in \mathcal{M}_{1}\right\}$ is a submanifold of $\mathcal{M}$, diffeomorphic to $\mathcal{M}_{1}$.
In particular if $m_{1}=1, R^{1}=0$, by Theorem 1.3 we obtain that stable critical points of the mapping $\left.S\right|_{\left(\mathcal{M}_{1}\right)\left(p_{2}\right)}$ yield existence of minimal discs inside (small) geodesic tubes around the curve $\left(\mathcal{M}_{1}\right)\left(p_{2}\right)$, where as before $\boldsymbol{S}$ is the scalar curvature of $\mathcal{M}$.
- As a simple byproduct of our analysis, we find that if $\Gamma$ is a closed curve, we have at least 2 (equal to the Lusternik-Schnierelman category of $\Gamma$, see [3]) solutions (without any assumptions on the curvature of $\mathcal{M}$ ).
- We believe that this result might be generalized to higher codimensions namely if $N^{\ell}, 1<$ $\ell<n$, is an $\ell$-dimensional submanifold of $\mathcal{M}^{n+1}$ and considering the following surface of revolution with axis of rotation $\mathbb{R}^{\ell}$

$$
S(s, z)=\left(\kappa^{1}(s), \ldots, \kappa^{\ell}(s), \phi(s) \Theta(z)\right)
$$

where $z \mapsto \Theta(z) \in S^{n-\ell}$, one could obtain $(n-\ell+1)$-dimensional minimal disc-type submanifolds of $\mathcal{M}$ centered on $N^{\ell}$.

Let us describe the proof of the theorems above. We first recall, see [17], that Capillary hyper-surfaces with constant contact angle arccos $\phi^{\prime}(s)$ are stationary for the energy functional

$$
\begin{equation*}
\mathcal{E}(D)=\operatorname{Area}\left(D \cap \Omega_{\rho}\right)-\phi^{\prime}(s) \operatorname{Area}\left(\Omega_{\rho}^{\prime}\right), \tag{3}
\end{equation*}
$$

among (orientable smooth) surfaces $D \subset \Omega_{\rho}$ with $\partial D \subset \partial \Omega_{\rho}$ and $\Omega_{\rho}^{\prime} \subset \partial \Omega_{\rho}$ is the part (on one side of $D$ ) for which the angle is measured. Moreover the Euler-Lagrange equations is nothing but

$$
\begin{array}{cccc}
H_{D} & = & 0 & \text { in } D, \\
\left\langle N_{D}, N_{\partial \Omega_{\rho}}\right\rangle & = & \phi^{\prime}(s) & \text { on } \partial D . \tag{4}
\end{array}
$$

Here $H_{D}$ is the mean curvature of $D$ while $N_{D}$ and $N_{\partial \Omega_{\rho}}$ are outer unit normals of $D$ and $\partial \Omega_{\rho}$ respectively. Since we look for stationary surfaces with a given profile for this energy functional, clearly by (1)-(2) a manifold of approximate solutions is given by $Z_{\rho}:=\left\{D_{s, \rho} \quad: \quad s \in[a, b]\right\}$. For any given hyper-surface $D_{s, \rho} \in Z_{\rho}$, we parametrize (locally) a neighborhood of $D_{s, \rho}$ (in the manifold in $\mathcal{M}$ ) by a mapping $F^{s}: \mathbb{R} \times B_{1}^{n} \rightarrow \mathcal{M}$ for which $F^{s}\left(t, \partial B_{1}^{n}\right) \subset \partial \Omega_{\rho}$, for every $t$, while the direction $F_{*}^{s}\left(\partial_{t}\right)$ is nearly normal to $D_{s, \rho}$, and moreover $D_{s, \rho}=F^{s}\left(0, B_{1}^{n}\right)$, see (8). This allows to parametrize any set $\mathscr{D}$ nearby $D_{s, \rho}$ satisfying $\partial \mathscr{D} \subset \partial \Omega_{\rho}$ by a function $w: B_{1}^{n} \rightarrow \mathbb{R}$ such that $\mathscr{D}(w)=F^{s}\left(w, B_{1}^{n}\right)$. We call $\mathcal{H}(s, \rho, w)$ the mean curvature of $\mathscr{D}(w)$ and $\mathcal{B}(s, \rho, w)$ the angle between the normals $N_{\partial \mathscr{D}(w)}$ and $N_{\mathscr{C} \rho}$ of $\partial \mathscr{D}(w)$ and $\partial \Omega_{\rho}$ respectively.
One of the main features in this work in the (technical) Sections 4.1, $\S 4.2$ is to calculate $\mathcal{H}(s, \rho, w)$ as a nonlinear elliptic partial differential operator, depending on $\rho$ and $s$ acting on $w$ coupled with the mixed boundary operator which we denote by $\mathcal{B}(s, \rho, w)$. In these calculations it is important to gather various different types of error terms, some of which depend linearly and some nonlinearly on $w$, and some of which are inhomogeneous terms vanishing to some order in $\rho$. It turns out to be helpful to rescale the local coordinates $y$ by $\varepsilon(s)=\rho \phi(s)$ which is the radius of the discs. The final expression, Proposition 4.5, for the mean curvature of $\mathscr{D}(w)$ then is

$$
-\frac{\phi}{\kappa^{\prime}} \mathcal{H}(s, \rho, w)=-\mathbb{L}_{\rho, s}(w)+\mathcal{O}\left(\rho^{2}\right)+\rho Q(w) \quad \text { in } \mathscr{D}(w)
$$

where $\mathbb{L}_{\rho, s}$ is the linearized mean curvature operator about $\mathscr{D}(0)=D_{s, \rho}$ :

$$
\mathbb{L}_{\rho, s}(w)=-\Delta w+\rho L_{s}(w) \quad \text { in } \mathscr{D}(w) ;
$$

also the angle between the normals satisfies (see Proposition 4.6)

$$
\rho^{-1}\left(\mathcal{B}(s, \rho, w)-\phi^{\prime}(s)\right)=\mathbb{B}_{\rho, s}(w)+\mathcal{O}(1)+\rho \bar{Q}(w)
$$

where

$$
\mathbb{B}_{\rho, s}(w)=\left(\left(\kappa^{\prime}(s)\right)^{2} \frac{\partial w}{\partial \eta}+\phi \phi^{\prime \prime} w\right)+\rho \bar{L}_{s}(w) \quad \text { on } \partial \mathscr{D}(w) .
$$

Here $L_{s}$ (resp. $\bar{L}_{s}$ ) is a second order (resp. first order) differential operator and $Q(w), \bar{Q}(w)$ are quadratic in $w$, see also the end of Section 2 for more precise definitions.

It turns out that the problem of finding $w$ such that $\mathscr{D}(w)$ solves (4) namely

$$
\begin{aligned}
\mathcal{H}(s, \rho, w) & =0 \\
\mathcal{B}(s, \rho, w) & =\phi^{\prime}(s)
\end{aligned} \quad \text { on } \mathscr{D}(w), ~(w), ~ l
$$

can be transformed to a fixed point problem for which the solvability is based on the invertibility of $\mathbb{L}_{\rho, s}$ on a suitable space of functions $w$ such that $\mathbb{B}_{\rho, s}(w)=0$. If $\phi \phi^{\prime \prime}>0$, the operator $\mathbb{L}_{\rho, s}$ (resp. $-\mathbb{L}_{\rho, s}$ ) is invertible by means of usual Sobolev inequalities. Hence after suitable adjustment of the disc $\mathscr{D}(w)$, we readily prove the first theorem. This program is carried out in § 5.1.

Now in the situation where $\phi \equiv 1$ and $\kappa=\mathrm{Id}$, it is clear that the linearized mean curvature $\mathbb{L}_{\rho, s}$ about any $D \in Z_{\rho}$ may have small (possibly zero) eigenvalues on the space of functions for which $\mathbb{B}_{\rho, s}(w)=\frac{\partial w}{\partial \eta}+\rho \bar{L}_{s}(w)=0$. This is related to the invariance by translations along the axis of rotation in the "flat" case. Hence $\mathbb{L}_{\rho, s}$ may not be invertible on such space. However restricting again ourselves on space of function orthogonal to the constant function 1 , we can perturb $Z_{\rho}$ to a manifold $\mathcal{Z}_{\rho}$ (constituted by sets having constant and small mean curvatures, see $\S$ 5.3.1) which turns out to be a natural constraint for $\mathcal{E}$ namely critical point of $\left.\mathcal{E}\right|_{\mathcal{Z}_{\rho}}$ is also stationary for $\mathcal{E}$. For that we use an argument from Kapouleas in [11] which was successfully employed in [15]. We will follow the argument of the latter, we refer to § 5.3.1.
It is worth noticing that this method is also closely related to variational-perturbative methods introduced by Ambrosetti and Badiale in [1] and subsequently used with success to get existence and multiplicity results for a wide class of variational problems in some perturbative setting we refer to the book by Ambrosetti Malchiodi [2] for more details and related applications.

Remark 1.5 - If $\phi^{\prime \prime} \equiv 0$, namely when $\phi(s)=c s+d$, $S$ parameterizes the cone we cannot conclude. However we notice that the proof of Theorem 1.3 highlights that near a point $\gamma\left(\rho s_{0}\right)$ for which $\phi^{\prime \prime}\left(s_{0}\right)=0$, there will be capillary surfaces with constant and small mean curvatures, see Remark 5.1.

- An interesting question is also to perturb the set

$$
\exp _{\gamma(\rho \kappa(s))}\left(\rho \phi(s) S^{n-1}\right)
$$

to a closed minimal submanifold of $\mathscr{C}^{\rho}$. One can see also the work by S.Secchi [19].

- Another problem can be set as follows. Let $\mathcal{U}$ be a smooth bounded domain of $\mathcal{M}$, and $\Gamma \hookrightarrow \partial \mathcal{U}$ be a smooth curve. We let $\tilde{y}=\left(y^{1}, \ldots, y^{n}\right)$ and $N_{\partial \mathcal{U}}$ be a unit interior normal field along $\partial \mathcal{U}$. Choosing an oriented orthogonal frame $\left(E_{1} \ldots, E_{n-1}\right)$ along $\Gamma$ in $\partial \mathcal{U}$, one obtains a coordinate system by letting, for any $y=\left(\tilde{y}, y^{n}\right)=\left(y^{1}, \ldots, y^{n-1}, y^{n}\right)$,

$$
F\left(x_{0}, \tilde{y}, y^{n}\right):=\exp _{\exp _{\gamma\left(x_{0}\right)}^{\mathcal{M}}\left(y^{i} E_{i}\right)}^{\mathcal{M}}\left(y^{n} N_{\partial \mathcal{U}}\right)
$$

Now consider the set

$$
F\left(\rho \kappa(s), \rho \phi(s) B_{+}^{n}\right),
$$

where $B_{+}^{n}=\left\{x=\left(x^{1}, \ldots, x^{n-1}, x^{n}\right) \in \mathbb{R}^{n} \quad: \quad|x|=1, \quad x^{n}>0\right\}$. One may be tempted to perturb the set above into capillary minimal surfaces that meet the "half"-surface of revolution

$$
F\left(\rho \kappa(s), \rho \phi(s) S_{+}^{n-1}\right)
$$

by an angle equal to arccos $\phi^{\prime}$. In this case, as we believe, a result like Theorem 1.1 would carry over. On the other hand one would need maybe to impose some conditions on the principal curvature of $\partial \mathcal{U}$ along $\Gamma$ in order to obtain a variant of Theorem 1.3.

## Acknowledgments

The authors would like to thank Prof. Antonio Ambrosetti for taking their attention to these questions. They are grateful to Prof. Antonio Ambrosetti and Prof. Andrea Malchiodi at SISSA/ISAS for the helpful discussions. They have been supported by M.U.R.S.T within the PRIN 2008 Variational Methods and Nonlinear Differential Equations.

## 2 Preliminaries and notations

We consider $(\phi, \kappa):[a, b] \rightarrow \mathbb{R}^{2}$ smooth with $\kappa^{\prime}(s), \phi(s) \neq 0$ for every $s \in[a, b]$ moreover we assume that $s$ is the arc length of the rotating curve $s \mapsto(\phi(s), \kappa(s))$ precisely

$$
\left(\phi^{\prime}(s)\right)^{2}+\left(\kappa^{\prime}(s)\right)^{2}=1 \quad \forall s \in[a, b] .
$$

We also assume that $x_{0}$ is the arc length of $\gamma$, and we will let $E_{0}:=\gamma^{\prime}$. We choose a parallel (local) orthonormal frame $E_{1}, \cdots, E_{n}$ of $N \Gamma$ along $\Gamma$. This determines a coordinate system by defining

$$
f\left(x_{0}, y\right):=\exp _{\gamma\left(x_{0}\right)}\left(y^{i} E_{i}\right) \quad \text { for } y=\left(y^{1}, \cdots, y^{n}\right)
$$

which therefore defines coordinates vector fields:

$$
Y_{0}:=f_{*}\left(\partial_{x_{0}}\right), \quad Y_{i}:=f_{*}\left(\partial_{y_{i}}\right)
$$

We will adopt the convention that the indices $i, j, k, \cdots \in\{1, \ldots, n\}$ while $\alpha, \beta, \cdots \in\{0, \ldots, n\}$ with $Y_{\alpha}=Y_{0}$ when $\alpha=0$.
By construction, $\left.\nabla_{X_{i}} Y_{0}\right|_{\Gamma} \in T \Gamma$ so that we can define

$$
\left.\left\langle\nabla_{X_{i}} Y_{0}, Y_{0}\right\rangle\right|_{\Gamma}=-\Gamma_{0}^{0}\left(E_{i}\right)
$$

There also holds

$$
\begin{equation*}
\nabla_{Y_{i}} Y_{j}(y)=O(|y|)_{\gamma} Y_{\gamma} . \tag{5}
\end{equation*}
$$

If $q=f\left(x_{0}, y\right) \in \mathcal{M}$ near the point $p=f\left(x_{0}, 0\right) \in \Gamma$, we can expand the metric $g_{\alpha \beta}(q)=$ $\left\langle Y_{\alpha}, Y_{\beta}\right\rangle$ in $y$, (we refer to [13] Proposition 2.1 for the proof).

Lemma 2.1 In the above coordinates $\left(x_{0}, y\right)$, for any $i, j=1, \ldots, n$, we have

$$
\begin{aligned}
g_{i j}(q) & =\delta_{i j}+\frac{1}{3}\left\langle R_{p}\left(Y, E_{i}\right) Y, E_{j}\right\rangle+\frac{1}{6}\left\langle\nabla_{Y} R_{p}\left(Y, E_{i}\right) Y, E_{j}\right\rangle+O_{p}\left(|y|^{4}\right) \\
g_{0 j}(q) & =\frac{2}{3}\left\langle R_{p}\left(Y, E_{0}\right) Y, E_{j}\right\rangle+O_{p}\left(|y|^{3}\right) \\
g_{00}(q) & =1-2 \Gamma_{0}^{0}(Y)+\left\langle R_{p}\left(Y, E_{0}\right) Y, E_{0}\right\rangle+O_{p}\left(|y|^{3}\right)
\end{aligned}
$$

where $Y:=y^{i} E_{i}$.
Notation for error terms: Any expression of the form $L(\omega)$ (resp. $\bar{L}(\omega)$ ) denotes a linear combination of the function $\omega$ together with its derivatives with respect to the vector fields $Y_{i}$ up to order 2 (resp. order 1 ). The coefficients of $L$ or $\bar{L}$ might depend on $\rho$ and $s$ but, for all $k \in \mathbb{N}$, there exists a constant $c>0$ independent of $\rho \in(0,1)$ and $s \in[a, b]$ such that

$$
\begin{gathered}
\|L(\omega)\|_{\mathcal{C}^{k, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c\|\omega\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)}, \\
\left\|\bar{L}_{s}(\omega)\right\|_{\mathcal{C}^{k, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c\|\omega\|_{\mathcal{C}^{k+1, \alpha}\left(\overline{B_{+}^{n}}\right)} .
\end{gathered}
$$

Similarly, any expression of the form $Q(\omega)$ (resp $\bar{Q}(\omega))$ denotes a nonlinear operator in the function $\omega$ together with its derivatives with respect to the vector fields $Y_{i}$ up to order 2 (resp. 1). The coefficients of the Taylor expansion of $Q^{a}(\omega)$ in powers of $\omega$ and its partial derivatives might depend on $\rho$ and $s$ and, given $k \in \mathbb{N}$, there exists a constant $c>0$ independent of $\rho \in(0,1)$ and $s \in[a, b]$ such that

$$
\left\|Q\left(\omega_{1}\right)-Q\left(\omega_{2}\right)\right\|_{\mathcal{C}^{k, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c\left(\left\|\omega_{1}\right\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)}+\left\|\omega_{2}\right\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)}\right)\left\|\omega_{1}-\omega_{2}\right\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)},
$$

provided $\left\|\omega_{i}\right\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq 1, i=1,2$. Also

$$
\left\|\bar{Q}\left(\omega_{1}\right)-\bar{Q}\left(\omega_{2}\right)\right\|_{\mathcal{C}^{k, \alpha}\left(\overline{B_{1}^{\bar{n}}}\right)} \leq c\left(\left\|\omega_{1}\right\|_{\mathcal{C}^{k+1, \alpha}\left(\overline{B_{1}^{n}}\right)}+\left\|\omega_{2}\right\|_{\mathcal{C}^{k+1, \alpha}\left(\overline{B_{1}^{n}}\right)}\right)\left\|\omega_{1}-\omega_{2}\right\|_{\mathcal{C}^{k+1, \alpha}\left(\overline{B_{1}^{n}}\right)},
$$

provided $\left\|\omega_{i}\right\|_{\mathcal{C}^{k+2, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq 1$. We also agree that any term denoted by $\mathcal{O}\left(r^{d}\right)$ ( with $r \in \mathbb{R}$ may depend on $s$ ) is a smooth function on $B_{1}^{n}$ that might depend on $s$ but satisfies

$$
\left\|\frac{\mathcal{O}\left(r^{d}\right)}{|r|^{d}}\right\|_{\mathcal{C}^{k, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c
$$

for a constant $c$ independent of $s$.

## 3 On the surface of revolution around $\Gamma$

We start by fixing the following notations which will be useful later.

## Notations:

Through the following of this paper,

$$
\varepsilon(s)=\rho \phi(s) \quad \text { and } \quad \varepsilon_{1}(s)=\rho \kappa(s) \quad \text { for every } s \in[a, b] .
$$

In terms of cylindrical coordinates, letting $\Theta(z): \mathbb{R}^{n-1} \rightarrow S^{n-1}$, the surface of revolution $\mathscr{C}^{\varepsilon}$ around $\Gamma$ can be parameterized by

$$
\mathcal{C}^{\varepsilon}(s, z):=f\left(\varepsilon_{1}(s), \varepsilon(s) \Theta(z)\right)=\exp _{\gamma\left(\varepsilon_{1}(s)\right)}\left(\varepsilon(s) \Theta^{i}(z) E_{i}\right)
$$

The tangent plane is spanned by the vector fields

$$
\begin{aligned}
& Z_{0}^{c}=\mathcal{C}^{\rho}\left(\partial_{x_{0}}\right)=\varepsilon_{1}^{\prime} Y_{0}+\varepsilon^{\prime} \Upsilon, \\
& Z_{j}^{c}=\mathcal{C}^{\rho}\left(\partial_{z^{j}}\right)=\varepsilon \Upsilon_{j}, \quad j=1, \cdots, n,
\end{aligned}
$$

where

$$
\Upsilon=\Theta^{i} Y_{i} .
$$

We recall also from [13]
Lemma 3.1 Let $q=\mathcal{C}^{\rho}(s, z) \in \mathscr{C}^{\rho}$, there hold

$$
\begin{aligned}
\langle\Upsilon, \Upsilon\rangle_{q} & =1 \\
\left\langle\Upsilon, Y_{0}\right\rangle_{q} & =0 \\
\left\langle\Upsilon, \Upsilon_{j}\right\rangle_{q} & =0 .
\end{aligned}
$$

Lemma 3.2 In the notations above, the first fundamental form of $\mathscr{C}^{\rho}$ has the following expansions

$$
\begin{aligned}
& \left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle=\frac{\varepsilon^{2}}{\phi^{2}}-2 \varepsilon\left|\varepsilon_{1}^{\prime}\right|^{2} \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left|\varepsilon_{1}^{\prime}\right|^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{5}\right), \\
& \left\langle Z_{l}^{c}, Z_{k}^{c}\right\rangle=\varepsilon^{2}\left\langle\Theta_{l}, \Theta_{k}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right), \\
& \left\langle Z_{0}^{c}, Z_{k}^{c}\right\rangle=\mathcal{O}\left(\varepsilon^{4}\right) .
\end{aligned}
$$

Proof. Recalling that

$$
\left|\varepsilon_{1}^{\prime}\right|^{2}+\left|\varepsilon^{\prime}\right|^{2}=\frac{\varepsilon^{2}}{\phi^{2}}
$$

hence we obtain, using also the Lemmas 3.1, 2.1, that

$$
\begin{aligned}
\left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle & =\left|\varepsilon_{1}^{\prime}\right|^{2}\left(1-2 \varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right)\right)+\left|\varepsilon^{\prime}\right|^{2} \\
& =\frac{\varepsilon^{2}}{\phi^{2}}-2 \varepsilon\left|\varepsilon_{1}^{\prime}\right|^{2} \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left|\varepsilon_{1}^{\prime}\right|^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{5}\right)
\end{aligned}
$$

The other expansions are easy consequences of the Lemmas 3.1 2.1.

### 3.1 The unit normal field to the surface of revolution

Call

$$
M:=\varepsilon^{\prime} X_{0}-\varepsilon_{1}^{\prime} \Upsilon
$$

and set

$$
\tilde{N}_{c}(s, z)=M+\alpha_{0} Z_{0}^{c}+\alpha_{k} Z_{k}^{c} .
$$

Note that this vector filed is normal (not necessary unitary) to the surface whenever we can determined $\alpha_{k}$ so that $\left\langle\tilde{N}_{c}, Z_{k}^{c}\right\rangle=\left\langle\tilde{N}_{c}, Z_{0}^{c}\right\rangle=0$ for all $k=1, \ldots, n$. This therefore leads to solving a linear system.

Observe that

$$
\begin{aligned}
\left\langle M, Z_{0}^{c}\right\rangle & =\varepsilon_{1}^{\prime} \varepsilon^{\prime}\left\langle Y_{0}, Y_{0}\right\rangle-\varepsilon_{1}^{\prime} \varepsilon^{\prime}\langle\Upsilon, \Upsilon\rangle \\
& =\varepsilon_{1}^{\prime} \varepsilon^{\prime}\left(1-2 \varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle-1+\mathcal{O}\left(\varepsilon^{3}\right)\right),
\end{aligned}
$$

hence

$$
\left\langle M, Z_{0}^{c}\right\rangle=-2 \varepsilon_{1}^{\prime} \varepsilon^{\prime} \varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon_{1}^{\prime} \varepsilon^{\prime} \varepsilon^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{5}\right) .
$$

Also we have

$$
\left\langle M, Z_{k}^{c}\right\rangle=\varepsilon \varepsilon^{\prime}\left\langle X_{0}, Y_{k}\right\rangle-\varepsilon_{1}^{\prime} \varepsilon\left\langle\Upsilon, Y_{k}\right\rangle=\mathcal{O}\left(\varepsilon^{4}\right) .
$$

If we use Lemma 3.2, we have

$$
\begin{aligned}
\alpha_{0}\left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle & =-\alpha_{k}\left\langle Z_{k}^{c}, Z_{0}^{c}\right\rangle-\left\langle M, Z_{0}^{c}\right\rangle \\
& =\alpha_{k} \mathcal{O}\left(\varepsilon^{4}\right)+2 \varepsilon \varepsilon_{1}^{\prime} \varepsilon^{\prime} \Gamma_{0}^{0}(\Theta)-\varepsilon^{2} \varepsilon_{1}^{\prime} \varepsilon^{\prime}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{5}\right)
\end{aligned}
$$

so
(6) $\alpha_{0}=\frac{\varepsilon^{\prime} \varepsilon_{1}^{\prime}}{\varepsilon^{2}} \phi^{2}\left(2 \varepsilon \Gamma_{0}^{0}(\Theta)-\left|\varepsilon_{1}^{\prime}\right|^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle\right)+4 \phi^{2} \frac{\left|\varepsilon_{1}^{\prime}\right|^{3}}{\varepsilon^{2}} \varepsilon^{\prime} \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)+\mathcal{O}\left(\varepsilon^{3}\right)+\alpha_{k} \mathcal{O}\left(\varepsilon^{4}\right)$.

Since

$$
\alpha_{k}\left\langle Z_{k}^{c}, Z_{l}^{c}\right\rangle+\alpha_{0}\left\langle Z_{0}^{c}, Z_{l}^{c}\right\rangle=-\left\langle M, Z_{l}^{c}\right\rangle
$$

and using (6)

$$
\alpha_{k}\left\langle Z_{k}^{c}, Z_{l}^{c}\right\rangle+\alpha_{k} \mathcal{O}\left(\varepsilon^{6}\right)+\mathcal{O}\left(\varepsilon^{5}\right)=\mathcal{O}\left(\varepsilon^{4}\right)
$$

we get

$$
\alpha_{k}\left(\varepsilon^{2}\left\langle\Theta_{l}, \Theta_{k}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right)\right)=\mathcal{O}\left(\varepsilon^{4}\right)
$$

therefore

$$
\alpha_{k}=\mathcal{O}\left(\varepsilon^{2}\right) .
$$

Recalling that $\varepsilon=\rho \phi$ while $\varepsilon_{1}=\rho \kappa$ we define $\bar{\alpha}_{0}$ by the relation

$$
\alpha_{0}=\phi^{\prime} \bar{\alpha}_{0}+\mathcal{O}\left(\varepsilon^{3}\right)
$$

Namely

$$
\bar{\alpha}_{0}=2 \varepsilon_{1}^{\prime} \phi \Gamma_{0}^{0}(\Theta)-\varepsilon^{\prime} \varepsilon_{1}^{\prime} \phi\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+4 \varepsilon_{1}^{\prime} \kappa^{\prime} \varepsilon \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta) .
$$

Now let us compute the norm of this normal vector field. Since

$$
\tilde{N}_{c}(s, z):=M+\alpha_{0} Z_{0}^{c}+\alpha_{k} Z_{k}^{c}
$$

we have by construction

$$
\left\langle\tilde{N}_{c}, \tilde{N}_{c}\right\rangle=\langle M, M\rangle+a_{0}^{2}\left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle+\alpha_{k} \alpha_{l}\left\langle Z_{k}^{c}, Z_{l}^{c}\right\rangle+2 \alpha_{0}\left\langle M, Z_{0}^{c}\right\rangle+2 \alpha_{k}\left\langle M, Z_{k}^{c}\right\rangle+2 \alpha_{k} \alpha_{0}\left\langle Z_{k}^{c}, Z_{0}^{c}\right\rangle .
$$

Notice that

$$
\begin{aligned}
\alpha_{0}\left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle & =-\left\langle M, Z_{0}^{c}\right\rangle-\alpha_{k}\left\langle Z_{k}^{c}, Z_{0}^{c}\right\rangle \\
& =-\left\langle M, Z_{0}^{c}\right\rangle+\mathcal{O}\left(\varepsilon^{6}\right)
\end{aligned}
$$

and

$$
\alpha_{0}^{2}\left\langle Z_{0}^{c}, Z_{0}^{c}\right\rangle=-\alpha_{0}\left\langle M, Z_{0}^{c}\right\rangle+\mathcal{O}\left(\varepsilon^{6}\right),
$$

hence

$$
\left\langle\tilde{N}_{c}, \tilde{N}_{c}\right\rangle=\langle M, M\rangle+\alpha_{0}\left\langle M, Z_{0}^{c}\right\rangle+\mathcal{O}\left(\varepsilon^{6}\right) .
$$

Now observe that

$$
\begin{aligned}
\langle M, M\rangle & =\left|\varepsilon^{\prime}\right|^{2}\left\langle X_{0}, X_{0}\right\rangle+\left|\varepsilon_{1}^{\prime}\right|^{2}\langle\Upsilon, \Upsilon\rangle \\
& =\frac{\varepsilon^{2}}{\phi^{2}}+\left|\varepsilon^{\prime}\right|^{2}\left(-2 \varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right)\right)
\end{aligned}
$$

and

$$
\alpha_{0}\left\langle M, Z_{0}^{c}\right\rangle=-2 \varepsilon \varepsilon_{1}^{\prime} \varepsilon^{\prime} \Gamma_{0}^{0}(\Theta) \alpha_{0}+\mathcal{O}\left(\varepsilon^{4}\right) \alpha_{0}=-4 \frac{\varepsilon^{2}}{\rho^{2}}\left|\varepsilon^{\prime}\right|^{2}\left|\varepsilon_{1}^{\prime}\right|^{2} \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)+\mathcal{O}\left(\varepsilon^{5}\right) .
$$

So we have

$$
\frac{\phi^{2}}{\varepsilon^{2}}\left\langle\tilde{N}_{c}, \tilde{N}_{c}\right\rangle=1+\frac{\left|\varepsilon^{\prime}\right|^{2}}{\varepsilon^{2}} \phi^{2}\left(-2 \varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon^{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle\right)-4 \frac{\phi^{4}}{\varepsilon^{2}}\left|\varepsilon^{\prime}\right|^{2}\left|\varepsilon_{1}^{\prime}\right|^{2} \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)+\mathcal{O}\left(\varepsilon^{3}\right)
$$

Finally we conclude that

$$
\frac{\varepsilon}{\phi}\left|\tilde{N}_{c}\right|^{-1}=1+\frac{\left|\varepsilon^{\prime}\right|^{2}}{\varepsilon} \phi^{2} \Gamma_{0}^{0}(\Theta)+\left(3\left|\frac{\varepsilon^{\prime}}{\varepsilon}\right|^{4} \phi^{4} \varepsilon^{2}+2\left|\frac{\varepsilon_{1}^{\prime}}{\varepsilon}\right|^{2}\left|\varepsilon^{\prime}\right|^{2} \phi^{4}\right) \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)-\frac{\left|\varepsilon^{\prime}\right|^{2}}{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right)
$$

and, setting

$$
H_{c}(\Theta, \Theta):=\left(3 \frac{\left|\varepsilon^{\prime}\right|^{2}}{\varepsilon^{2}}+2 \frac{\left|\varepsilon_{1}^{\prime}\right|^{2}}{\varepsilon^{2}}\right) \phi^{2} \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)-\frac{1}{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle
$$

we can simply write

$$
\frac{\varepsilon}{\phi}\left|\tilde{N}_{c}\right|^{-1}=1+\left|\varepsilon^{\prime}\right|^{2} \phi^{2}\left[\frac{1}{\varepsilon} \Gamma_{0}^{0}(\Theta)+H_{c}(\Theta, \Theta)\right]+\mathcal{O}\left(\varepsilon^{3}\right)
$$

We collect all these in the following

Proposition 3.3 There exists an interior (non unit) normal vector field of $\mathscr{C}^{\rho}$ which has the following expansions

$$
\tilde{N}_{\mathscr{C} P}(s, z)=-\varepsilon_{1}^{\prime} \Upsilon+\varepsilon^{\prime} Y_{0}+\left(\phi^{\prime} \bar{\alpha}_{0}+\mathcal{O}\left(\varepsilon^{3}\right)\right) Z_{0}^{c}+\alpha_{k} Z_{k}^{c},
$$

where

$$
\begin{gathered}
\bar{\alpha}_{0}=2 \varepsilon_{1}^{\prime} \phi \Gamma_{0}^{0}(\Theta)-\varepsilon^{\prime} \varepsilon_{1}^{\prime} \phi\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+4 \varepsilon_{1}^{\prime} \kappa^{\prime} \varepsilon \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta) ; \\
\alpha_{k}=\mathcal{O}\left(\varepsilon^{2}\right) .
\end{gathered}
$$

Moreover

$$
\rho\left|\tilde{N}_{\mathscr{C} \rho}\right|^{-1}=1+\left|\phi^{\prime}\right|^{2}\left(\varepsilon \Gamma_{0}^{0}(\Theta)+\varepsilon^{2} H_{c}(\Theta, \Theta)\right)+\mathcal{O}\left(\varepsilon^{3}\right),
$$

where

$$
H_{c}(\Theta, \Theta)=\left(\left|\phi^{\prime}\right|^{2}+2 \phi^{4}\right) \Gamma_{0}^{0}(\Theta) \Gamma_{0}^{0}(\Theta)-\frac{1}{2}\left\langle R\left(\Theta, E_{0}\right) \Theta, E_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right)
$$

## 4 Discs centered on $\Gamma$ with boundary on $\mathscr{C}^{\rho}$

For $\delta>0, B_{\delta}^{n}$ will denote the ball of $\mathbb{R}^{n}$ with radius $\delta$ centered at the origin. For any $s$, we consider the disc $\mathscr{D}^{s, \varepsilon}$ of radius $\varepsilon$ centered at $\gamma(\varepsilon(s))$ given by

$$
\mathscr{D}^{s, \varepsilon}:=f\left(\varepsilon_{1}(s), \varepsilon(s) B_{1}^{n}\right),
$$

parameterized by

$$
B_{1}^{n} \ni x \mapsto \mathcal{D}^{s, \varepsilon}(x)=f\left(\varepsilon_{1}, \varepsilon x\right) .
$$

## Notations

$$
\bar{\varepsilon}(s, t):=\varepsilon(s+\varepsilon(s) t) \quad \bar{\varepsilon}_{1}(s, t):=\varepsilon_{1}(s+\varepsilon(s) t) ;
$$

$\bar{\varepsilon}^{\prime}(s, t):=\partial_{t} \varepsilon(s+\varepsilon(s) t)=\varepsilon(s) \varepsilon^{\prime}(s+\varepsilon(s) t) \quad \bar{\varepsilon}_{1}^{\prime}(s, t):=\partial_{t} \varepsilon_{1}(s+\varepsilon(s) t)=\varepsilon(s) \varepsilon_{1}^{\prime}(s+\varepsilon(s) t)$.

Notice that

$$
\begin{equation*}
\bar{\varepsilon}^{\prime}(s, t)=\left(\varepsilon^{\prime}+\varepsilon \varepsilon^{\prime \prime} t+\varepsilon^{3} \mathcal{O}\left(t^{2}\right)\right) \varepsilon, \quad \bar{\varepsilon}_{1}^{\prime}(s, t)=\left(\varepsilon_{1}^{\prime}+\varepsilon \varepsilon_{1}^{\prime \prime} t+\varepsilon^{3} \mathcal{O}\left(t^{2}\right)\right) \varepsilon \tag{7}
\end{equation*}
$$

A parametrization of the neighborhood of the disc (in $\mathcal{M})$ centered at $p=\gamma\left(\varepsilon_{1}(s)\right) \in \Gamma$ with radius $\varepsilon$ ( $\rho$ small) can be defined by

$$
\begin{equation*}
F^{s}(t, x):=f\left(\bar{\varepsilon}_{1}(s, t), \bar{\varepsilon}(s, t) x\right) \quad \forall x \in B_{1}^{n},|t| \ll 1 \tag{8}
\end{equation*}
$$

note that by construction,

$$
F^{s}\left(t, \partial B_{1}^{n}\right) \subset \mathscr{C}^{\rho}, \quad \forall|t| \ll 1
$$

and more precisely, for every $|t| \ll 1$

$$
F^{s}(t, x)=\mathcal{C}^{\rho}(s+\varepsilon t, \varepsilon(s+\varepsilon t) x) \quad \forall x \in \partial B_{1}^{n}=S^{n-1}
$$

We consider the following vector fields induced by $F^{s}$

$$
\begin{aligned}
\varepsilon T_{0} & :=F_{*}\left(\partial_{t}\right)=\bar{\varepsilon}_{1}^{\prime} Y_{0}(t)+\bar{\varepsilon}^{\prime} X(t), \\
\bar{\varepsilon} T_{j} & :=F_{*}\left(\partial_{y^{j}}\right)=\bar{\varepsilon} Y_{j}(t) .
\end{aligned}
$$

Here $X=x^{i} E_{i}$.

Lemma 4.1 At $q=F^{s}(t, x)$, we have

$$
\begin{aligned}
\left\langle Y_{i}(t), Y_{j}(t)\right\rangle_{q} & =\delta_{i j}+\frac{\varepsilon^{2}}{3}\left\langle R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\frac{\varepsilon^{3}}{6}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\mathcal{O}_{p}\left(\varepsilon^{4}\right)+\varepsilon^{3} \mathcal{O}(t)+\varepsilon^{4} \mathcal{O}_{p}\left(t^{2}\right) ; \\
\left\langle Y_{0}(t), Y_{j}(t)\right\rangle_{q} & =\frac{2 \varepsilon^{2}}{3}\left\langle R_{p}\left(X, E_{0}\right) X, E_{j}\right\rangle+\mathcal{O}_{p}\left(\varepsilon^{3}\right)+\varepsilon^{3} \mathcal{O}_{p}(t)+\varepsilon^{4} \mathcal{O}_{p}\left(t^{2}\right) \\
\left\langle Y_{0}(t), Y_{0}(t)\right\rangle_{q} & =1-2 \bar{\varepsilon} \Gamma_{0}^{0}(X)+2 \varepsilon^{2} U_{0}^{0}(X) t+\varepsilon^{2}\left\langle R_{p}\left(X, E_{0}\right) X, E_{0}\right\rangle+\mathcal{O}_{p}\left(\varepsilon^{3}\right)+\varepsilon^{3} \mathcal{O}_{p}(t)+\varepsilon^{4} \mathcal{O}_{p}\left(t^{2}\right),
\end{aligned}
$$

where $p=\gamma\left(\varepsilon_{1}(s)\right) \in \Gamma$ and

$$
\varepsilon U_{0}^{0}(X)=\varepsilon_{1}^{\prime} \Gamma_{00}^{0}-\varepsilon^{\prime} \Gamma_{0}^{0}(X)
$$

Proof. There holds

$$
\begin{aligned}
\left.\frac{d}{d t}\left\langle Y_{0}(t), Y_{0}(t)\right\rangle_{q}\right|_{t=0} & =\left.2\left\langle\nabla_{\varepsilon T_{0}} Y_{0}, Y_{0}\right\rangle\right|_{t=0} \\
& =\left.2 \bar{\varepsilon}_{1}^{\prime}\left\langle\nabla_{Y_{0}} Y_{0}, Y_{0}\right\rangle\right|_{t=0}+\left.2 \bar{\varepsilon}^{\prime}\left\langle\nabla_{X} Y_{0}, Y_{0}\right\rangle\right|_{t=0} \\
& =2 \varepsilon\left(\varepsilon_{1}^{\prime} \Gamma_{00}^{0}-\varepsilon^{\prime} \Gamma_{0}^{0}(X)\right)+\mathcal{O}_{p}\left(\varepsilon^{3}\right)
\end{aligned}
$$

we have used (7). Hence the last expansion follows. On the other hand one has

$$
\begin{aligned}
\left.\frac{d}{d t}\left\langle Y_{0}(t), Y_{i}(t)\right\rangle_{q}\right|_{t=0} & =\left.\left\langle\nabla_{\varepsilon T_{0}} Y_{0}, Y_{i}\right\rangle\right|_{t=0}+\left.\left\langle Y_{0}, \nabla_{\varepsilon T_{0}} Y_{i}\right\rangle\right|_{t=0} \\
& =\left.\bar{\varepsilon}_{1}^{\prime}\left\langle\nabla_{Y_{0}} Y_{0}, Y_{i}\right\rangle\right|_{t=0}+\left.2 \bar{\varepsilon}^{\prime}\left\langle\nabla_{X} Y_{0}, Y_{i}\right\rangle\right|_{t=0}+\left.\bar{\varepsilon}_{1}^{\prime}\left\langle Y_{0}, \nabla_{Y_{0}} Y_{i}\right\rangle\right|_{t=0}+\left.2 \bar{\varepsilon}^{\prime}\left\langle Y_{0}, \nabla_{X} Y_{i}\right\rangle\right|_{t=0}
\end{aligned}
$$

Since by construction $\left\langle\nabla_{E_{0}} E_{0}, E_{i}\right\rangle+\left\langle E_{0}, \nabla_{E_{0}} E_{i}\right\rangle=0$ and also $\left\langle\nabla_{X} E_{0}, E_{i}\right\rangle+\left\langle E_{0}, \nabla_{X} E_{i}\right\rangle=0$ on $\Gamma$, we infer that

$$
\left.\frac{d}{d t}\left\langle Y_{0}(t), Y_{i}(t)\right\rangle_{q}\right|_{t=0}=\mathcal{O}_{p}\left(\varepsilon^{3}\right)
$$

In the same vain, the first expansions follows similarly.

Using the above lemma, and (7) we get
Lemma 4.2 The following expansions hold

$$
\begin{aligned}
\left\langle T_{i}, T_{j}\right\rangle & =\delta_{i j}+\frac{\varepsilon^{2}}{3}\left\langle R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\frac{\varepsilon^{3}}{6}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\mathcal{O}_{p}\left(\varepsilon^{4}\right)+\varepsilon^{3} \mathcal{O}(t)+\varepsilon^{4} \mathcal{O}_{p}\left(t^{2}\right) ; \\
\left\langle T_{0}, T_{j}\right\rangle & =\varepsilon^{\prime} x^{i}+\mathcal{O}_{p}\left(\varepsilon^{3}\right)+\varepsilon^{4} \mathcal{O}_{p}(t)+\varepsilon^{5} \mathcal{O}_{p}\left(t^{2}\right) ; \\
\left\langle T_{0}, T_{0}\right\rangle & =\frac{\varepsilon^{2}}{\phi^{2}}+\left|\varepsilon_{1}^{\prime}\right|^{2}\left(-2 \bar{\varepsilon} \Gamma_{0}^{0}(X)+2 \varepsilon^{2} U_{0}^{0}(X) t\right)+\mathcal{O}_{p}\left(\varepsilon^{5}\right)+\varepsilon^{5} \mathcal{O}_{p}(t)+\varepsilon^{6} \mathcal{O}_{p}\left(t^{2}\right),
\end{aligned}
$$

where $p=\gamma\left(\varepsilon_{1}(s)\right) \in \Gamma$.
Observe that all disc-type surfaces nearby $\mathscr{D}_{s, \rho}$ with boundary contained in $\mathscr{C}^{\rho}$ can be parameterized by

$$
\begin{equation*}
G^{s}(x):=F^{s}(w(x), x), \tag{9}
\end{equation*}
$$

for some smooth function $w: B_{1}^{n} \rightarrow \mathbb{R}$. We will call $\mathscr{D}^{s, \varepsilon(s)}(w)=G^{s}\left(B_{1}^{n}\right)$.

### 4.1 Mean curvature of Perturbed disc $\mathscr{D}(w)$

It is not difficult to see that the tangent plane of $\mathscr{D}(w)=\mathscr{D}^{s, \varepsilon(s)}(w)$ is spanned by the vector fields

$$
Z_{j}=G_{*}^{s}\left(\partial_{x^{j}}\right)=\varepsilon w_{x^{j}} T_{0}+\bar{\varepsilon}(s, w) T_{j} .
$$

From Lemma 4.2, it is clear that at the point $q=G^{s}(x)=F^{s}(w(x), x)$ there hold (10)

$$
\begin{aligned}
\left\langle T_{i}, T_{j}\right\rangle_{q} & =\delta_{i j}+\frac{\varepsilon^{2}}{3}\left\langle R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\frac{\varepsilon^{3}}{6}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{3} L(w)+\varepsilon^{4} Q(w) \\
\left\langle T_{0}, T_{j}\right\rangle_{q} & =\varepsilon^{\prime} x^{j}+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{4} L(w)+\varepsilon^{5} Q(w) \\
\left\langle T_{0}, T_{0}\right\rangle_{q} & =\frac{\varepsilon^{2}}{\phi^{2}}+\left|\varepsilon_{1}^{\prime}\right|^{2}\left(-2 \bar{\varepsilon} \Gamma_{0}^{0}(X)+2 \varepsilon^{2} U_{0}^{0}(X) w\right)+\mathcal{O}_{p}\left(\varepsilon^{5}\right)+\varepsilon^{5} L(w)+\varepsilon^{6} Q(w)
\end{aligned}
$$

Observing that $\bar{\varepsilon}(s, w)=\varepsilon+\varepsilon \varepsilon^{\prime} w+\varepsilon^{3} Q(w)$ and using (10), we get the first fundamental form $h_{i j}:=\left\langle Z_{i}, Z_{j}\right\rangle$,

$$
\begin{aligned}
\varepsilon^{-2} h_{i j} & =\left(1+2 \varepsilon^{\prime} w\right) \delta_{i j}+\varepsilon^{\prime}\left(w_{x^{i}} x^{j}+w_{x^{j}} x^{i}\right)+\frac{\varepsilon^{2}}{3}\left\langle R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\frac{\varepsilon^{3}}{6}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle \\
& +\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

### 4.1.1 The normal vector field

Considering the vector field

$$
\tilde{N}_{\mathscr{D}}=Y_{0}+a_{k} Z_{k}
$$

Observe that it is normal (not necessary unitary) to the disc whenever we can find $a_{k}$ such that $\left\langle\tilde{N}_{\mathscr{D}}, Z_{k}\right\rangle=0$ for any $k=1, \ldots, n$. Namely $a_{k}$ satisfies

$$
\begin{equation*}
\mathscr{D}^{s, \varepsilon(s)}(w) a_{k} h_{i k}=-\left\langle Y_{0}, Z_{i}\right\rangle . \tag{12}
\end{equation*}
$$

Since

$$
\left\langle Y_{0}, Z_{i}\right\rangle=\bar{\varepsilon}_{1}^{\prime} w_{x^{i}}\left\langle Y_{0}, Y_{0}\right\rangle+\bar{\varepsilon}^{\prime} w_{x^{i}}\left\langle Y_{0}, X\right\rangle+\bar{\varepsilon}\left\langle Y_{0}, Y_{i}\right\rangle
$$

then from (11) and (7), we get the formula

$$
\begin{equation*}
\varepsilon^{2} a_{k}=\varepsilon \varepsilon_{1}^{\prime} w_{x^{k}}\left\langle Y_{0}, Y_{0}\right\rangle+\varepsilon\left\langle Y_{0}, Y_{k}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{4} L(w)+\varepsilon^{3} Q(w) \tag{13}
\end{equation*}
$$

And also since $\bar{\varepsilon}=\varepsilon+\varepsilon^{2} L(w), \bar{\varepsilon}_{1}=\varepsilon_{1}+\varepsilon^{2} L(w)$, we get

$$
\varepsilon^{2} a_{k}=-\varepsilon \varepsilon_{1}^{\prime}\left(1-2 \varepsilon \Gamma_{0}^{0}(X)\right) w_{x^{k}}-\frac{2 \varepsilon^{3}}{3}\left\langle R\left(X, E_{0}\right) X, E_{k}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{4} L(w)+\varepsilon^{3} Q(w)
$$

and thus

$$
a_{k}=-\frac{\varepsilon_{1}^{\prime}}{\varepsilon}\left(1-2 \varepsilon \Gamma_{0}^{0}(X)\right) w_{x^{i}}-\frac{2 \varepsilon}{3}\left\langle R\left(X, E_{0}\right) X, E_{i}\right\rangle+\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon Q(w)
$$

Moreover using also (12) we have

$$
\begin{aligned}
\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{D}}\right\rangle & =\left\langle Y_{0}, Y_{0}\right\rangle-a_{k} a_{l} h_{k l} \\
& =\left\langle Y_{0}, Y_{0}\right\rangle-a_{k}\left(\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{2} L(w)+\varepsilon^{4} Q(w)\right) \\
& =\left\langle Y_{0}, Y_{0}\right\rangle-\left(\mathcal{O}(\varepsilon)+L(w)+\varepsilon^{2} Q(w)\right)\left(\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{2} L(w)+\varepsilon^{2} Q(w)\right) \\
& =\left\langle Y_{0}, Y_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

Hence

$$
\begin{equation*}
\left|\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{D}}\right\rangle\right|^{-1}=\left|\left\langle Y_{0}, Y_{0}\right\rangle\right|^{-1}+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) \tag{14}
\end{equation*}
$$

Therefore

$$
\begin{equation*}
\left|\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{D}}\right\rangle\right|^{-1}=1+\bar{\varepsilon} \Gamma_{0}^{0}(X)+\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon^{2} Q(w) \tag{15}
\end{equation*}
$$

We then conclude that the unit normal has the following expansions:

$$
\begin{align*}
N_{\mathscr{D}} & =\left(1+\varepsilon \Gamma_{0}^{0}(X)+\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon^{2} Q(w)\right) Y_{0} \\
& +\left(-\frac{\varepsilon_{1}^{\prime}}{\varepsilon}\left(1+\varepsilon \Gamma_{0}^{0}(X)\right) w_{x^{k}}-\frac{2 \varepsilon}{3}\left\langle R\left(X, E_{0}\right) X, E_{k}\right\rangle+\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon Q(w)\right) Z_{k} . \tag{16}
\end{align*}
$$

Sometimes we will simply need to write $N_{\mathscr{D}}$ in the more compact form

$$
N_{\mathscr{D}}=Y_{0}+\left(\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon L(w)+\varepsilon^{2} Q(w)\right)_{\alpha} Y_{\alpha} .
$$

### 4.1.2 The Second Fundamental Form

Observe that in the scaled variables $y=\varepsilon x$, since the functions $\mathcal{O}\left(\varepsilon^{m}\right), L(w)$ and $Q(w)$ are depending on $x$ whereas the vector fields $Y_{\alpha}$ depend on $y=\varepsilon x$, we have for any

$$
E_{i}\left(\mathcal{O}\left(\varepsilon^{m}\right)\right)=\mathcal{O}\left(\varepsilon^{m-1}\right), \quad E_{i}\left(\varepsilon^{m} L(w)\right)=\varepsilon^{m-1} L(w), \quad E_{i}\left(\varepsilon^{m} Q(w)\right)=\varepsilon^{m-1} Q(w)
$$

Having this in mind, we state the following
Lemma 4.3 There holds

$$
\left\langle T_{0}, \nabla_{Z_{i}} N_{\mathscr{D}}\right\rangle=\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon L(w)+\varepsilon^{2} Q(w) .
$$

Proof. Using (16) and recall that $T_{0}=\varepsilon_{1}^{\prime} Y_{0}+\varepsilon^{\prime} X+\varepsilon L(w)_{\alpha} Y_{\alpha}$, we have

$$
\begin{aligned}
\left.\left\langle T_{0}, \nabla_{Z_{i}} N_{\mathscr{D}}\right\rangle\right|_{w=0} & =\left\langle T_{0}, \nabla_{\varepsilon Y_{i}}\left(1+\varepsilon \Gamma_{0}^{0}(X)\right) Y_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right) \\
& =\varepsilon \Gamma_{0}^{0}\left(E_{i}\right)\left\langle T_{0}, Y_{0}\right\rangle+\varepsilon\left(1+\varepsilon \Gamma_{0}^{0}(X)\right)\left\langle T_{0}, \nabla_{Y_{i}} Y_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right) \\
& =\varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{i}\right)\left\langle Y_{0}, Y_{0}\right\rangle+\varepsilon \varepsilon_{1}^{\prime}\left(1+\varepsilon \Gamma_{0}^{0}(X)\right)\left\langle Y_{0}, \nabla_{Y_{i}} Y_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{3}\right) \\
& =\varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{i}\right)-\varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{i}\right)+\mathcal{O}\left(\varepsilon^{3}\right) .
\end{aligned}
$$

Hence we get the result.
Let us now estimate the second fundamental form of $\mathscr{D}(w)$.

## Lemma 4.4

$$
\begin{aligned}
\left\langle\nabla_{Z_{i}} Z_{j}, N_{\mathscr{D}}\right\rangle & =\varepsilon\left(1-\varepsilon \Gamma_{0}^{0}\left(E_{l}\right) x^{l}\right) w_{x^{i} x^{j}}+\varepsilon^{2}\left\langle\nabla_{Y i} Y_{j}, Y_{0}\right\rangle+\mathcal{O}\left(\varepsilon^{4}\right) \\
& -\varepsilon^{2}\left(w_{x^{j}} \Gamma_{0}^{0}\left(E_{i}\right)+w_{x^{i}} \Gamma_{0}^{0}\left(E_{j}\right)\right)+\varepsilon^{3} L(w)+\varepsilon^{3} Q(w)
\end{aligned}
$$

Proof. we have

$$
\left\langle\nabla_{Z_{i}} Z_{j}, N_{\mathscr{D}}\right\rangle=\varepsilon\left\langle\nabla_{Z_{i}}\left(w_{x^{j}} T_{0}\right), N\right\rangle+\left\langle\nabla_{Z_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle .
$$

We first estimate $\left\langle\nabla_{Z_{i}}\left(w_{x^{j}} T_{0}\right), N_{\mathscr{D}}\right\rangle$.
Observe that

$$
\frac{\partial}{\partial x^{i}}\left\langle w_{x^{j}} T_{0}, N_{\mathscr{D}}\right\rangle=\left\langle\nabla_{Z_{i}}\left(w_{x^{j}} T_{0}\right), N_{\mathscr{D}}\right\rangle+\left\langle w_{x^{j}} T_{0}, \nabla_{Z_{i}} N_{\mathscr{D}}\right\rangle
$$

which implies that

$$
\left\langle\nabla_{Z_{i}}\left(w_{x^{j}} T_{0}\right), N_{\mathscr{D}}\right\rangle=\frac{\partial}{\partial x^{i}}\left\langle w_{x^{j}} T_{0}, N_{\mathscr{D}}\right\rangle-w_{x^{j}}\left\langle T_{0}, \nabla_{Z_{i}} N_{\mathscr{D}}\right\rangle .
$$

The formula (12) shows that

$$
\left\langle Y_{0}, \tilde{N}_{\mathscr{D}}\right\rangle=\left\langle Y_{0}, Y_{0}\right\rangle+a_{k}\left\langle Z_{k}, Y_{0}\right\rangle=\left\langle Y_{0}, Y_{0}\right\rangle-a_{k} a_{l}\left\langle Z_{k}, Z_{l}\right\rangle=\left|\tilde{N}_{\mathscr{D}}\right|^{2}
$$

and then

$$
\left\langle Y_{0}, N_{\mathscr{D}}\right\rangle=\left|\tilde{N}_{\mathscr{D}}\right| .
$$

From the fact that $\left\langle Y_{0}, X\right\rangle=0$ when $w=0$ and that

$$
\left\langle Z_{k}, X\right\rangle=\varepsilon x^{k}+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{2} L(w)+\varepsilon^{5} Q(w)
$$

we obtain $a_{k}\left\langle Z_{k}, X\right\rangle=\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon L(w)+\varepsilon^{2} Q(w)$, from which the following hold

$$
\begin{aligned}
\left\langle\varepsilon T_{0}, N_{\mathscr{D}}\right\rangle & =\bar{\varepsilon}_{1}^{\prime}\left\langle Y_{0}, N_{\mathscr{D}}\right\rangle+\bar{\varepsilon}^{\prime}\left\langle X, N_{\mathscr{D}}\right\rangle \\
& =\bar{\varepsilon}_{1}^{\prime}\left|\tilde{N}_{\mathscr{D}}\right|+\bar{\varepsilon}^{\prime}\left(\varepsilon^{2}+\varepsilon L(w)+\varepsilon^{2} Q(w)\right) \\
& =\varepsilon \varepsilon_{1}^{\prime}\left(1-\varepsilon \Gamma_{0}^{0}(X)\right)+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{3} L(w)+\varepsilon^{4} Q(w) .
\end{aligned}
$$

From this, we deduce that

$$
\frac{\partial}{\partial x^{i}}\left\langle w_{x^{j}} T_{0}, N_{\mathscr{D}}\right\rangle=\varepsilon_{1}^{\prime}\left(1-\varepsilon \Gamma_{0}^{0}(X)\right) w_{x^{i} x^{j}}-\varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{i}\right) w_{x^{j}}+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
$$

We conclude using also Lemma 4.3 that

$$
\begin{equation*}
\left\langle\nabla_{Z_{i}}\left(w_{x^{j}} T_{0}\right), N_{\mathscr{D}}\right\rangle=\varepsilon_{1}^{\prime}\left(1-\varepsilon \Gamma_{0}^{0}(X)\right) w_{x^{i} x^{j}}-\varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{i}\right) w_{x^{j}}+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) . \tag{17}
\end{equation*}
$$

It remains the term $\left\langle\nabla_{Z_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle=\varepsilon w_{x^{i}}\left\langle\nabla_{T_{0}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle+\bar{\varepsilon}\left\langle\nabla_{T_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle$.
Since

$$
\begin{equation*}
\bar{\varepsilon}(s, w)=\varepsilon(s)+\varepsilon L(w)+\varepsilon^{3} Q(w) \tag{18}
\end{equation*}
$$

we can write

$$
\left\langle\nabla_{T_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle=\varepsilon\left\langle\nabla_{T_{i}} T_{j}, N\right\rangle+\left\langle\nabla_{T_{i}}\left(\left(\varepsilon^{2} L+\varepsilon^{3} Q\right) T_{j}\right), N_{\mathscr{D}}\right\rangle .
$$

Recalling that $\nabla_{Y_{i}} Y_{j}=\left(\mathcal{O}(\varepsilon)+\varepsilon L+\varepsilon^{2} Q\right)_{\alpha} Y_{\alpha}$ also $\left\langle T_{i}, N_{\mathscr{D}}\right\rangle=\varepsilon^{2}+\varepsilon L+\varepsilon^{2} Q$ thus

$$
\left\langle\nabla_{T_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle=\left.\varepsilon\left\langle\nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle\right|_{w=0}+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{3} Q(w)
$$

Moreover (7) and (18) yield

$$
\left.\left\langle\nabla_{T_{0}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle\right|_{w=0}=\varepsilon \varepsilon \varepsilon_{1}^{\prime}\left\langle\nabla_{Y_{0}} Y_{j}, N_{\mathscr{D}}\right\rangle+\varepsilon \varepsilon \varepsilon^{\prime}\left\langle\nabla_{X} Y_{j}, N_{\mathscr{D}}\right\rangle=-\varepsilon \varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}+\mathcal{O}\left(\varepsilon^{4}\right) .
$$

This implies that

$$
\left\langle\nabla_{T_{0}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle=-\varepsilon \varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}+\mathcal{O}\left(\varepsilon^{4}\right)+\varepsilon^{2} L(w)+\varepsilon^{3} Q(w) .
$$

Finally, collecting these and using (18) it turns out that

$$
\begin{equation*}
\left\langle\nabla_{Z_{i}}\left(\bar{\varepsilon} T_{j}\right), N_{\mathscr{D}}\right\rangle=\left.\varepsilon \varepsilon\left\langle\nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle\right|_{w=0}+\mathcal{O}\left(\varepsilon^{4}\right)-\varepsilon \varepsilon \varepsilon_{1}^{\prime} \Gamma_{0}^{0}\left(E_{j}\right) w_{x^{i}}+\varepsilon^{4} L(w)+\varepsilon^{3} Q(w) \tag{19}
\end{equation*}
$$

The result follows from (17) and (19).

We need also to expand more precisely $\left.\left\langle\nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle\right|_{w=0}$. By construction it vanish on $\Gamma$ and

$$
Y_{l}\left\langle\nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle=\left\langle\nabla_{Y_{l}} \nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle+\left\langle\nabla_{Y_{i}} Y_{j}, \nabla_{Y_{l}} Y_{0}\right\rangle
$$

Furthermore by (5) and since (see for instance [8] Lemma 9.20)

$$
\left.\nabla_{Y_{l}} \nabla_{Y_{i}} Y_{j}\right|_{\gamma\left(x^{0}\right)}=-\frac{1}{3}\left(R\left(E_{l}, E_{i}\right) E_{j}+R\left(E_{l}, E_{j}\right) E_{i}\right)
$$

it follows that

$$
\left\langle\nabla_{Y_{i}} Y_{j}, Y_{0}\right\rangle=-\frac{\varepsilon}{3}\left(\left\langle R\left(X, E_{i}\right) E_{j}, E_{0}\right\rangle+\left\langle R\left(X, E_{j}\right) E_{i}, E_{0}\right\rangle\right)+\mathcal{O}\left(\varepsilon^{2}\right)
$$

We conclude that from Lemma 4.4 that the Second fundamental form $\amalg_{i j}=\left\langle\nabla_{Z_{i}} Z_{j}, N^{\mathscr{D}}\right\rangle$ of the perturbed disc $\mathscr{D}^{s, \varepsilon(s)}(w)$ centered at the point $\gamma(\varepsilon(s))$ with radius $\varepsilon(s)$ is given by

$$
\begin{align*}
\amalg_{i j}= & \varepsilon \varepsilon_{1}^{\prime}\left(1-\varepsilon \Gamma_{0}^{0}\left(E_{l}\right) x^{l}\right) w_{x^{j} x^{i}}-\frac{\varepsilon^{3}}{3}\left(\left\langle R\left(X, E_{i}\right) E_{j}, E_{0}\right\rangle+\left\langle R\left(X, E_{j}\right) E_{i}, E_{0}\right\rangle\right)  \tag{20}\\
& +\mathcal{O}\left(\varepsilon^{4}\right)-\varepsilon^{2} \varepsilon_{1}^{\prime}\left(w_{x^{j}} \Gamma_{0}^{0}\left(E_{i}\right)+w_{x^{i}} \Gamma_{0}^{0}\left(E_{j}\right)\right)+\varepsilon^{4} L(w)+\varepsilon^{3} Q(w) .
\end{align*}
$$

We recall that if $E_{\alpha}$ is an orthogonal basis of $T_{p} \mathcal{M}$, then

$$
\operatorname{Ric}_{p}(X, Y)=-\left\langle R_{p}\left(X, E_{\alpha}\right) Y, E_{\alpha}\right\rangle \quad \forall X, Y \in T_{p} \mathcal{M}
$$

Finally we obtain

$$
\begin{aligned}
\amalg_{i j} h^{i j}= & \frac{\varepsilon_{1}^{\prime}}{\varepsilon}\left(1-\varepsilon \Gamma_{0}^{0}(X)\right) \Delta w-\frac{2 \varepsilon}{3} \operatorname{Ric}_{p}\left(X, E_{0}\right) \\
& +\mathcal{O}\left(\varepsilon^{2}\right)-2 \varepsilon_{1}^{\prime} \Gamma_{0}^{0}(\nabla w)+\varepsilon^{2} L(w)+\varepsilon Q(w),
\end{aligned}
$$

where $X=x^{l} E_{l}$.

Proposition 4.5 In the above notations, the mean curvature $\mathcal{H}(s, \rho, w)$ of $\mathscr{D}_{s, \rho}(w)$ has the following expansions

$$
\frac{\phi}{\kappa^{\prime}} \mathcal{H}(s, \rho, w)=\Delta w-\frac{2 \rho}{3} \frac{\phi^{2}}{\kappa^{\prime}} \operatorname{Ric}_{p}\left(X, E_{0}\right)+\mathcal{O}\left(\rho^{2}\right)+\rho L(w)+\rho Q(w) .
$$

In particular if $\Gamma$ is a geodesic, $\Gamma_{0}^{0}=0$ then

$$
\frac{\phi}{\kappa^{\prime}} \mathcal{H}(s, \rho, w)=\Delta w-\frac{2 \rho}{3} \frac{\phi^{2}}{\kappa^{\prime}} \operatorname{Ric}_{p}\left(X, E_{0}\right)+\mathcal{O}\left(\rho^{2}\right)+\rho^{2} L(w)+\rho Q(w)
$$

### 4.2 Angle between the normals

By construction, at $q=G^{s}(x)$ we have $F^{s}(w(x), x)=\mathcal{C}^{\varepsilon}(s+\varepsilon w(x), \varepsilon(s+\varepsilon w(x)))$, for every $x \in \partial B_{1}^{n}$. Recall from § 3.1 and 4.1.1 that

$$
\tilde{N}_{\mathscr{C} \rho}(s, z)=-\varepsilon_{1}^{\prime}(s) X+\varepsilon^{\prime}(s) Y_{0}+\alpha_{0} Z_{0}^{c}+\alpha_{k} Z_{k}^{c},
$$

where $\alpha_{0}=\phi^{\prime} \bar{\alpha}_{0}+\mathcal{O}\left(\varepsilon^{3}\right)$ and $\alpha_{k}=\mathcal{O}\left(\varepsilon^{2}\right)$ also

$$
\tilde{N}_{\mathscr{D}}=Y_{0}+a_{k} Z_{k} .
$$

One easily verifies that

$$
\begin{aligned}
\left\langle\tilde{N}_{\mathscr{D}}, \rho^{-1} \tilde{N}_{\mathscr{C}}(s+\varepsilon w)\right\rangle_{q} & =-\kappa^{\prime}\left\langle X, Y_{0}\right\rangle_{q}+\phi^{\prime}\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\frac{\alpha_{0}}{\rho}\left\langle Z_{0}^{c}, Y_{0}\right\rangle_{q}+\frac{\alpha_{k}}{\rho}\left\langle Z_{k}^{c}, Y_{0}\right\rangle_{q} \\
& -\kappa^{\prime} a_{k}\left\langle X, Z_{k}\right\rangle_{q}+\phi^{\prime} a_{k}\left\langle Z_{k}, Y_{0}\right\rangle_{q}+\frac{a_{k}}{\rho} \alpha_{0}\left\langle Z_{0}^{c}, Z_{k}\right\rangle_{q}+\frac{\alpha_{k}}{\rho} a_{l}\left\langle Z_{k}, Z_{l}^{c}\right\rangle_{q} .
\end{aligned}
$$

We have to expand

$$
\kappa^{\prime}(s+\varepsilon w)=\kappa^{\prime}(s)+\varepsilon \kappa^{\prime \prime}(s) w+\varepsilon^{2} Q(w) \quad \phi^{\prime}(s+\varepsilon w)=\phi^{\prime}(s)+\varepsilon \phi^{\prime \prime}(s) w+\varepsilon^{2} Q(w) .
$$

We will also need the following result which uses just the expansions of the metric Lemma 4.1

$$
\begin{align*}
\left\langle Z_{0}^{c}, Z_{k}\right\rangle & =\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon^{3} Q(w) \\
\left\langle Z_{l}^{c}, Z_{k}\right\rangle & =\mathcal{O}\left(\varepsilon^{2}\right)+\varepsilon^{2} L(w)+\varepsilon^{3} Q(w)  \tag{21}\\
\left\langle X, Z_{k}\right\rangle_{q} & =\left(\varepsilon+\varepsilon^{\prime} \varepsilon w\right) x^{k}+\varepsilon^{3} L(w)+\varepsilon^{5} Q(w) .
\end{align*}
$$

We use the fact that $a_{k}\left\langle Z_{l}, Z_{k}\right\rangle=-\left\langle Y_{0}, Z_{k}\right\rangle$ to have

$$
\begin{aligned}
\rho^{-1}\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{C} \varepsilon}(s+\varepsilon w)\right\rangle_{q} & =\left(\phi^{\prime}+\varepsilon w \phi^{\prime \prime}\right)\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\left(\kappa^{\prime}+\varepsilon \kappa^{\prime \prime} w\right) \alpha_{0}\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\frac{\alpha_{k}}{\rho}\left\langle Z_{k}^{c}, Y_{0}\right\rangle_{q} \\
& -\left(\kappa^{\prime}+\varepsilon \kappa^{\prime \prime} w\right) a_{k}\left\langle X, Z_{k}\right\rangle_{q}-\phi^{\prime} a_{k} a_{l}\left\langle Z_{k}, Z_{l}\right\rangle_{q}+\frac{a_{k}}{\rho} \alpha_{0}\left\langle Z_{0}^{c}, Z_{k}\right\rangle_{q}+\frac{\alpha_{k}}{\rho} a_{l}\left\langle Z_{k}, Z_{l}^{c}\right\rangle_{q} \\
& +\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

Now from (21) we get

$$
-\phi^{\prime} a_{k} a_{l}\left\langle Z_{k}, Z_{l}\right\rangle_{q}+\frac{\alpha_{0}}{\rho} a_{k}\left\langle Z_{0}^{c}, Z_{k}\right\rangle_{q}+\frac{\alpha_{k}}{\rho} a_{l}\left\langle Z_{k}, Z_{l}^{c}\right\rangle_{q}=\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w)
$$

and also since

$$
\left\langle Z_{k}^{c}, Y_{0}\right\rangle_{q}=\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{4} Q(w)
$$

one has

$$
\begin{aligned}
\rho^{-1}\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{C} \varepsilon}(s+\varepsilon w)\right\rangle_{q} & =\left(\phi^{\prime}+\varepsilon w \phi^{\prime \prime}\right)\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\left(\kappa^{\prime}+\varepsilon w \kappa^{\prime \prime}\right) \bar{\alpha}_{0}\left\langle Y_{0}, Y_{0}\right\rangle_{q} \\
& -\left(\kappa^{\prime}+\varepsilon \kappa^{\prime \prime} w\right) a_{k}\left\langle X, Z_{k}\right\rangle_{q}+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

From (21) and recalling the formula for $a_{k}$ in (13) we get

$$
\begin{aligned}
a_{k}\left\langle X, Z_{k}\right\rangle_{q} & =-\varepsilon_{1}^{\prime} \frac{\partial w}{\partial \eta}\left\langle Y_{0}, Y_{0}\right\rangle_{q}-\left(1+\varepsilon^{\prime} w\right)\left\langle Y_{0}, X\right\rangle_{q}+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) \\
& =-\varepsilon_{1}^{\prime} \frac{\partial w}{\partial \eta}\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w)
\end{aligned}
$$

and then we deduce that

$$
\begin{aligned}
\rho^{-1}\left\langle\tilde{N}_{\mathscr{D}}, \tilde{N}_{\mathscr{C}}(s+\varepsilon w)\right\rangle_{q} & =\phi^{\prime}\left(1+\kappa^{\prime} \bar{\alpha}_{0}\right)\left\langle Y_{0}, Y_{0}\right\rangle_{q}+\left(\varepsilon w \phi^{\prime \prime}+\kappa^{\prime} \varepsilon_{1}^{\prime}\right) \frac{\partial w}{\partial \eta}\left\langle Y_{0}, Y_{0}\right\rangle_{q}+2 \varepsilon^{2} \kappa^{\prime \prime} \kappa^{\prime} w \Gamma_{0}^{0}(X) \\
& +\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

Using (14), we have that

$$
\begin{aligned}
\rho^{-1}\left\langle N_{\mathscr{D}}, \tilde{N}_{\mathscr{C}} \varepsilon(s+\varepsilon w)\right\rangle_{q} & =\phi^{\prime}\left(1+\kappa^{\prime} \bar{\alpha}_{0}\right)\left|Y_{0}\right|_{q}+\left(\varepsilon w \phi^{\prime \prime}+\kappa^{\prime} \varepsilon_{1}^{\prime}\right) \frac{\partial w}{\partial \eta}\left|Y_{0}\right|_{q}+2 \varepsilon^{2} \kappa^{\prime \prime} \kappa^{\prime} w \Gamma_{0}^{0}(X) \\
& +\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

Since $\alpha_{0}=\phi^{\prime}(s+\varepsilon w) \bar{\alpha}_{0}+\mathcal{O}\left(\varepsilon^{3}\right)$, one has

$$
\alpha_{0}=\left(\phi^{\prime}(s)+\phi^{\prime \prime}(s) \varepsilon(s) w\right) \bar{\alpha}_{0}+\varepsilon^{2} Q(w) \bar{\alpha}_{0}
$$

so that

$$
\alpha_{0}=\phi^{\prime}(s) \bar{\alpha}_{0}-2 \phi^{\prime \prime}(s) \varepsilon_{1}^{\prime} \varepsilon w \Gamma_{0}^{0}(X)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
$$

Moreover notice that

$$
\left.\Gamma_{0}^{0}(X)\right|_{q}=\Gamma_{0}^{0}(X)+\varepsilon^{2} L(w)+\varepsilon^{3} Q(w)
$$

and also

$$
\left|\phi^{\prime}(s+\varepsilon w)\right|^{2}=\left(\phi^{\prime}+2 \varepsilon w \phi^{\prime \prime}\right) \phi^{\prime}+\varepsilon^{2} Q(w),
$$

we have that
$\rho\left|\tilde{N}_{\mathscr{C} \rho}(s+\varepsilon w)\right|^{-1}=1+\left(\phi^{\prime}+2 \varepsilon \phi^{\prime \prime} w+\varepsilon^{\prime} w\right) \varepsilon \phi^{\prime} \Gamma_{0}^{0}(X)+\left|\phi^{\prime}(s)\right|^{2} \varepsilon^{2} H_{c}(X, X)+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{3} Q(w)$
from which we deduce that

$$
\begin{aligned}
\rho\left|\tilde{N}_{\mathscr{C} \rho}(s+\varepsilon w)\right|_{q}^{-1}\left|Y_{0}\right|_{q} & =1-\left(\kappa^{\prime}\right)^{2} \varepsilon \Gamma_{0}^{0}(X)+\varepsilon^{2}\left(3-\left(\phi^{\prime}\right)^{2}+\left(\phi^{\prime}\right)^{4}+2\left(\phi^{\prime}\right)^{2} \phi^{4}\right) \Gamma_{0}^{0}(X) \Gamma_{0}^{0}(X) \\
& +\frac{2-\left(\phi^{\prime}\right)^{2}}{2}\left\langle R_{p}\left(X, E_{0}\right) X, E_{0}\right\rangle+\varepsilon w\left(-\varepsilon^{\prime}+2 \varepsilon \phi \phi^{\prime \prime}+\varepsilon^{\prime} \phi^{\prime}\right) \Gamma_{0}^{0}(X)+2 \varepsilon^{2} U_{0}^{0}(X) w \\
& +\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w)
\end{aligned}
$$

Consequently we may expand the angle as

$$
\begin{aligned}
\left\langle N_{\mathscr{D}}, N_{\mathscr{C}}(s+\varepsilon w)\right\rangle_{q} & =\rho \phi^{\prime}\left(1+\kappa^{\prime} \bar{\alpha}_{0}\right)\left|Y_{0}\right|_{q}\left|\tilde{N}_{\mathscr{C} P}(s+\varepsilon w)\right|_{q}+2 \varepsilon^{2} \kappa^{\prime \prime} \kappa^{\prime} w \Gamma_{0}^{0}(X) \\
& +\left(1-\left(\kappa^{\prime}\right)^{2} \varepsilon \Gamma_{0}^{0}(X)\right)\left(\varepsilon w \phi^{\prime \prime}+\kappa^{\prime} \varepsilon_{1}^{\prime} \frac{\partial w}{\partial \eta}\right) \\
& +\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) \\
\left\langle N_{\mathscr{D}(w)}, N_{\mathscr{C} \varepsilon}(s+\varepsilon w)\right\rangle_{q}= & \phi^{\prime}(s)\left(1+\left(\kappa^{\prime}\right)^{2} \varepsilon \Gamma_{0}^{0}(X)\right)+\left(1-\left(\kappa^{\prime}\right)^{2} \varepsilon \Gamma_{0}^{0}(X)\right)\left(\varepsilon w \phi^{\prime \prime}+\kappa^{\prime} \varepsilon_{1}^{\prime} \frac{\partial w}{\partial \eta}\right) \\
& +\phi^{\prime} \varepsilon^{2}\left(3-2\left(\kappa^{\prime}\right)^{4}-\left(\phi^{\prime}\right)^{2}+\left(\phi^{\prime}\right)^{4}+2\left(\phi^{\prime}\right)^{2} \phi^{4}\right) \Gamma_{0}^{0}(X) \Gamma_{0}^{0}(X) \\
& +\frac{2-\left(\phi^{\prime}\right)^{2}}{2} \phi^{\prime} \varepsilon^{2}\left\langle R_{p}\left(X, E_{0}\right) X, E_{0}\right\rangle+\varepsilon \phi^{\prime} w\left(-\varepsilon^{\prime}+2 \varepsilon \kappa^{\prime \prime} \kappa^{\prime}+2 \varepsilon \phi \phi^{\prime \prime}+\varepsilon^{\prime} \phi^{\prime}\right) \Gamma_{0}^{0}(X) \\
& +2 \varepsilon^{2} U_{0}^{0}(X) w+\mathcal{O}\left(\varepsilon^{3}\right)+\varepsilon^{3} L(w)+\varepsilon^{2} Q(w) .
\end{aligned}
$$

We define

$$
\mathcal{B}(s, \rho, w):=\left\langle N_{\mathscr{D}(w)}, N_{\mathscr{C}} \varepsilon(s+\varepsilon w)\right\rangle_{q} .
$$

Now we conclude this section by collecting all these in the following
Proposition 4.6 In the above notations, there holds

$$
\begin{aligned}
\mathcal{B}(s, \rho, w) & =\phi^{\prime}(s)\left(1+\left(\kappa^{\prime}\right)^{2} \rho \phi \Gamma_{0}^{0}(X)\right)+\rho\left(\left(\kappa^{\prime}\right)^{2} \frac{\partial w}{\partial \eta}+\phi \phi^{\prime \prime} w\right) \\
& +\mathcal{O}\left(\rho^{2}\right)+\rho^{2} \bar{L}(w)+\rho^{2} \bar{Q}(w)
\end{aligned}
$$

while if $\phi^{\prime}(s)=0$, one has

$$
\mathcal{B}(s, \rho, w)=\rho\left(\kappa^{\prime}\right)^{2} \frac{\partial w}{\partial \eta}+\mathcal{O}\left(\rho^{3}\right)+\rho^{3} \bar{L}(w)+\rho^{2} \bar{Q}(w) .
$$

In particular if $\Gamma$ is a geodesic, we get precisely

$$
\begin{aligned}
\mathcal{B}(s, \rho, w) & =\phi^{\prime}(s)\left(1+\frac{1+\left(\kappa^{\prime}\right)^{2}}{2} \rho^{2} \phi^{2}\left\langle R_{p}\left(X, E_{0}\right) X, E_{0}\right\rangle\right)+\rho\left(\left(\kappa^{\prime}\right)^{2} \frac{\partial w}{\partial \eta}+\phi \phi^{\prime \prime} w\right) \\
& +\mathcal{O}\left(\rho^{3}\right)+\rho^{3} \bar{L}(w)+\rho^{2} \bar{Q}(w)
\end{aligned}
$$

## 5 Existence of capillary minimal submanifolds

### 5.1 Case where $\phi\left(s_{0}\right) \phi^{\prime \prime}\left(s_{0}\right)>0$

We may assume that $\phi(s) \phi^{\prime \prime}(s)>0$ for all $s \in I_{s_{0}}(\delta):=\left[s_{0}-\delta, s_{0}+\delta\right]$ for some $\delta>0$ small. We define the following operator by

$$
\left(\mathcal{L}_{s} w, v\right):=\int_{B_{1}^{n}} \nabla w \nabla v d x+\frac{\phi \phi^{\prime \prime}}{\left(\kappa^{\prime}\right)^{2}} \oint_{\partial B_{1}^{n}} w v d \sigma
$$

It is clear that from the inequality (see [22], Theorem A.9)

$$
\begin{equation*}
\int_{B_{1}^{n}} w^{2} d x \leq C(n)\left(\int_{B_{1}^{n}}|\nabla w|^{2} d x+\oint_{\partial B_{1}^{n}} w^{2} d \sigma\right), \quad \forall w \in H^{1} \tag{22}
\end{equation*}
$$

the operator $\mathcal{L}_{s}$ is coercive if $\rho$ is small. We call $w_{1}^{s, \rho}$ the unique solution to the equation

$$
\left(\mathcal{L}_{s} w_{1}, v\right):=-\phi \oint_{\partial B_{1}^{n}} \Gamma_{0}^{0}(X) v d \sigma
$$

Namely it solves the problem

$$
\begin{array}{cccc}
-\Delta w_{1} & = & 0 & \text { in } B_{1}^{n} \\
\frac{\partial w_{1}}{\partial \eta}+\frac{\phi \phi^{\prime \prime}}{\left(\kappa^{\prime}\right)^{2}} w_{1} & = & -\phi \Gamma_{0}^{0}(X) & \text { on } \partial B_{1}^{n}
\end{array}
$$

By elliptic regularity theory, there exist a constant $c>0$ (independent of $\rho$ and $s$ ) such that

$$
\left\|w_{1}^{s, \rho}\right\|_{\mathcal{C}^{2, \alpha}} \leq c \quad \forall s \in I_{s_{0}}(\delta)
$$

Moreover we have that for all $k \geq 0$

$$
\left\|\frac{\partial^{k} w_{1}^{s, \rho}}{\partial s^{k}}\right\|_{\mathcal{C}^{2, \alpha}} \leq c_{k} \quad \forall s \in I_{s_{0}}(\delta)
$$

for some constant $c_{k}$ which does not depend on $s$ nor on $\rho$ small.
Clearly by construction there holds

$$
\left\{\begin{array}{rlr}
\mathcal{H}\left(s, \rho, w_{1}^{s, \rho}\right) & =\mathcal{O}(\rho) & \text { in } \quad \mathscr{D}_{s, \rho}\left(w_{1}^{s, \rho}\right), \\
\mathcal{B}\left(s, \rho, w_{1}^{s, \rho}\right)=\phi^{\prime}(s)+\mathcal{O}\left(\rho^{2}\right) & \text { on } \quad \partial \mathscr{D}_{s, \rho}\left(w_{1}^{s, \rho}\right)
\end{array}\right.
$$

We define the space

$$
\begin{aligned}
\mathcal{C}_{s, \rho}^{2, \alpha} & :=\left\{w \in \mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right) \quad:\left.\quad \frac{\partial}{\partial v} \mathcal{B}\left(s, \rho, w_{1}^{s, \rho}+v\right)\right|_{v=0}[w]=0\right\} \\
& =\left\{w \in \mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right) \quad: \quad \frac{\partial w}{\partial \eta}+\frac{\phi \phi^{\prime \prime}}{\left(\kappa^{\prime}\right)^{2}} w+\rho \bar{L}_{s}(w)=0\right\}
\end{aligned}
$$

We consider the linearized mean curvature operator about $\mathscr{D}^{s, \rho}\left(w_{1}^{s, \rho}\right)$ (see Proposition 4.5), $\mathbb{L}_{\rho, s}(w): \mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right) \rightarrow \mathcal{C}^{0, \alpha}\left(\overline{B_{1}^{n}}\right)$ defined by

$$
\mathbb{L}_{\rho, s}(w):=-\left.\frac{\phi}{\kappa^{\prime}} \frac{\partial}{\partial v} \mathcal{H}\left(s, \rho, w_{1}^{s, \rho}+v\right)\right|_{v=0}[w]=-\Delta w+\rho L_{s}(w) .
$$

We define also $\Phi(s, \rho, x), \mathcal{Q}_{s, \rho}(w) \in \mathcal{C}^{0, \alpha}\left(\overline{B_{1}^{n}}\right)$ by duality as

$$
\left(\Phi(s, \rho, x), w^{\prime}\right):=-\frac{\phi}{\kappa^{\prime}} \int_{B_{1}^{n}} \mathcal{H}\left(s, \rho, w_{1}^{s, \rho}\right) w^{\prime} d x+\rho^{-1} \oint_{\partial B_{1}^{n}}\left(\mathcal{B}\left(s, \rho, w_{1}^{s, \rho}\right)-\phi^{\prime}(s)\right) w^{\prime} d s
$$

and for every $w \in \mathcal{C}^{2, \alpha}$

$$
\left(\mathcal{Q}_{s, \rho}(w), w^{\prime}\right):=\int_{B_{1}^{n}} Q(w) w^{\prime} d x+\oint_{\partial B_{1}^{n}} \bar{Q}(w) w^{\prime} d s, \quad \forall w^{\prime} \in L^{2}
$$

Clearly the solvability of the system

$$
\left\{\begin{array}{rlrr}
\mathcal{H}\left(s, \rho, w_{1}^{s, \rho}+w\right) & =0 & \text { in } & \mathscr{D}_{s, \rho}\left(w_{1}^{s, \rho}+w\right),  \tag{23}\\
\mathcal{B}\left(s, \rho, w_{1}^{s, \rho}+w\right) & =\phi^{\prime}(s) & \text { on } \quad \partial \mathscr{D}_{s, \rho}\left(w_{1}^{s, \rho}+w\right)
\end{array}\right.
$$

is equivalent to the fixed point problem

$$
\begin{equation*}
w=-\left(\left.\mathbb{L}_{\rho, s}\right|_{\mathcal{C}_{s, \rho}^{2, \alpha}}\right)^{-1}\left\{\Phi(s, \rho, x)+\rho \mathcal{Q}_{s, \rho}(w)\right\} \tag{24}
\end{equation*}
$$

Furthermore one has

$$
\begin{aligned}
\left\|\mathcal{Q}_{s, \rho}(w)\right\|_{\mathcal{C}^{0, \alpha}} & =O\left(\|w\|_{\mathcal{C}^{2, \alpha}}\right)\|w\|_{\mathcal{C}^{2, \alpha}}^{2} \\
\left\|\mathcal{Q}_{s, \rho}\left(w_{1}\right)-\mathcal{Q}_{s, \rho}\left(w_{2}\right)\right\|_{\mathcal{C}^{0, \alpha}} & =O\left(\left\|w_{1}\right\|_{\mathcal{C}^{2, \alpha}},\left\|w_{2}\right\|_{\mathcal{C}^{2, \alpha}}\right)\left\|w_{1}-w_{2}\right\|_{\mathcal{C}^{2, \alpha}}
\end{aligned}
$$

also by construction, there exist a constant $c>0$ (independent of $\rho$ and $s$ ) such that

$$
\|\Phi(s, \rho, \cdot)\|_{\mathcal{C}^{0, \alpha}} \leq c \rho \quad \forall s \in I_{s_{0}}(\delta)
$$

By (22) the operator $\mathbb{L}_{\rho, s}$ is coercive on $\mathcal{C}_{s, \rho}^{2, \alpha}$ if $\rho$ is small enough and also by elliptic regularity theory, $\mathbb{L}_{\rho, s}$ is an isomorphism from $\mathcal{C}_{s, \rho}^{2, \alpha}$ into $\mathcal{C}^{0, \alpha}\left(\overline{B_{1}^{n}}\right)$ therefore we can solve the fixed point problem (24) in a ball of $\mathcal{C}_{s, \rho}^{2, \alpha}$ with radius $C \rho$ for some $C>0$ which does not depend neither on $\rho$ small nor $s$.
And thus for $\rho$ small and $s \in I_{s_{0}}(\delta)$ there exists a function $w^{s, \rho} \in \mathcal{C}_{s, \rho}^{2, \alpha}$, with $\left\|w^{s, \rho}\right\|_{\mathcal{C}^{2, \alpha}} \leq C \rho$ such that

Namely $\mathscr{D}_{s, \rho}\left(w_{1}^{s, \rho}+w^{s, \rho}\right)$ is a capillary submanifold of $\Omega_{\rho}$ with constant contact angle $\arccos \phi^{\prime}(s)$ if $\rho$ is small enough by $\mathcal{C}^{2, \alpha}$ bound up to the boundary of $\tilde{w}^{s, \rho}=w_{1}^{s, \rho}+w^{s, \rho}$. Furthermore it follows from the construction that, for all $k \geq 0$

$$
\begin{equation*}
\left\|\frac{\partial^{k} \tilde{w}^{s, \rho}}{\partial s^{k}}\right\|_{\mathcal{C}^{2, \alpha}} \leq c_{k} \rho \quad \forall s \in I_{s_{0}}(\delta) \tag{25}
\end{equation*}
$$

for some constant $c_{k}$ which does not depend on $s$ nor on $\rho$ small.

### 5.2 Foliation by minimal discs

Call $\tilde{w}^{s, \rho}=w_{1}^{s, \rho}+w^{s, \rho}$. From (8), Lemma 4.1 and (25) the mapping

$$
I_{s_{0}}(\delta) \times B_{1}^{n} \ni(s, x) \xrightarrow{\Psi_{\rho}} F^{s}\left(\tilde{w}^{s, \rho}(x), x\right)=f\left(\bar{\varepsilon}_{1}\left(s, \tilde{w}^{s, \rho}(x)\right), \bar{\varepsilon}\left(s, \tilde{w}^{s, \rho}(x)\right) x\right)
$$

has Jacobian determinant which expands as

$$
\rho^{2 n+2}\left(\left(1-|x|^{2}\left(\phi^{\prime}\right)^{2}\right) \phi^{2 n}+\mathcal{O}_{s}(\rho)\right)
$$

and hence since $\left(\phi^{\prime}\right)^{2} \in(0,1)$ (see section 1 ), $\Psi_{\rho}$ is a local homeomorphism if $\rho$ is small enough. In particular it is a homeomorphism of a neighborhood of $\left(s_{0}, 0\right)$ which implies that there exist $0<\delta^{\prime}<\delta$ and $\varrho>0$ such that

$$
\Psi_{\rho}\left(s, B_{\varrho}^{n}\right) \cap \Psi_{\rho}\left(s^{\prime}, B_{\varrho}^{n}\right)=\emptyset \quad \forall s \neq s^{\prime} \in I_{s_{0}}\left(\delta^{\prime}\right)
$$

for every $\rho$ sufficiently small.
In this way the family of discs $\mathscr{D}_{s, \rho \varrho}\left(\tilde{w}^{s, \rho}\right), s \in I_{s_{0}}\left(\delta^{\prime}\right)$ with radius $\rho \varrho \phi(s)$ centered at $\gamma(\rho \kappa(s))$ constitutes a foliation of a neighborhood of $\gamma\left(\rho \kappa\left(s_{0}\right)\right)$ for which each leaf $\mathscr{D}_{s, \rho \varrho}\left(\tilde{w}^{s, \rho}\right)$ is a minimal disc intersecting $\mathscr{C}^{\rho \varrho}$ transversely along its boundary (the angle of contact may not be equal to $\left.\arccos \phi^{\prime}(s)\right)$.

## $5.3 \phi \equiv 1$ and $\kappa=\mathrm{Id}, \mathscr{C}^{\rho}$ is the geodesic tube around $\Gamma$

In this situation,

$$
\mathscr{C}^{\rho}=\left\{q \in \mathcal{M} \quad: \quad \operatorname{dist}_{g}(q, \Gamma)=\rho\right\}
$$

and its interior is

$$
\Omega_{\rho}=\left\{q \in \mathcal{M} \quad: \quad \operatorname{dist}_{g}(q, \Gamma)<\rho\right\}
$$

By [17], it is well known that (smooth) minimal surfaces $D \subset \Omega_{\rho}$ with $\partial D \subset \mathscr{C}^{\rho}$ are stationary for the area functional relative to $\mathscr{C}^{\rho}$ which is $D \mapsto \operatorname{Area}\left(D \cap \Omega_{\rho}\right)$ under variations $\Psi_{t}: D \rightarrow \mathcal{M}$ such that $\partial \Psi_{t}(D) \subset \mathscr{C}^{\rho}$ moreover the Euler-Lagrange equations are given by

$$
\begin{array}{ccc}
H_{D} & =0 & \text { in } D \\
\left\langle N_{D}, N_{\mathscr{C} P}\right\rangle & =0 & \text { on } \partial D . \tag{26}
\end{array}
$$

### 5.3.1 A finite-dimensional reduction

For every $s \in[a, b]$ and $X=x^{i} E_{i}$, we let $w_{1}^{s, \rho}$ be the solution of the following problem:

$$
\begin{aligned}
-\Delta w_{1} & =-\frac{2}{3} \operatorname{Ric}_{p}\left(X, E_{0}\right) & & \text { in } B_{1}^{n} \\
\frac{\partial w_{1}}{\partial \eta} & =0 & & \text { on } \partial B_{1}^{n}
\end{aligned}
$$

where $p=\gamma(\rho \kappa(s))$.
By elliptic regularity theory, there exist a constant $c>0$ (independent of $\rho$ and $s$ ) such that

$$
\begin{equation*}
\left\|w_{1}^{s, \rho}\right\|_{\mathcal{C}^{2, \alpha}} \leq c \quad \forall s \in[a, b] . \tag{27}
\end{equation*}
$$

As in $\S 5.1$, we let

$$
\begin{aligned}
\mathcal{C}_{s, \rho}^{2, \alpha} & :=\left\{w \in \mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right) \quad:\left.\quad \frac{\partial}{\partial v} \mathcal{B}\left(s, \rho, \rho w_{1}^{s, \rho}+v\right)\right|_{v=0}[w]=0\right\} \\
& =\left\{w \in \mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right) \quad: \quad \frac{\partial w}{\partial \eta}+\rho \bar{L}(w)=0\right\}
\end{aligned}
$$

As explained in the first section, the linearized mean curvature operator about $\mathscr{D}^{s, \rho}\left(\rho w_{1}^{s, \rho}\right)$ restricted on $\mathcal{C}_{s, \rho}^{2, \alpha}$ defined by

$$
\mathbb{L}_{\rho, s}(w):=-\left.\frac{\partial}{\partial v} \mathcal{H}\left(s, \rho, \rho w_{1}^{s, \rho}+v\right)\right|_{v=0}[w]=-\Delta w+\rho L_{s}(w)
$$

may have small (possibly zero) eigenvalues hence it may not be invertible on $\mathcal{C}_{s, \rho}^{2, \alpha}$. However instead of solving (26), we will prove that there exists a constant $\lambda_{s, \rho} \in \mathbb{R}$ and a function $w^{s, \rho} \in \mathcal{C}_{s, \rho}^{2, \alpha}$ such that

$$
\left\{\begin{array}{rlr}
\mathcal{H}\left(s, \rho, w^{s, \rho}\right)=\lambda_{s, \rho} & \text { in } \quad \mathscr{D}_{s, \rho}\left(w^{s, \rho}\right)  \tag{28}\\
\mathcal{B}\left(s, \rho, w^{s, \rho}\right)=0 & \text { on } \quad \partial \mathscr{D}_{s, \rho}\left(w^{s, \rho}\right)
\end{array}\right.
$$

To achieve this we let $P$ be the $L^{2}$ projection on the space of functions $w \in L^{2}$ which are orthogonal to the constant function $1, \int_{B_{1}^{n}} w d x=0$. Now if $\rho$ is small enough, the Poincare inequality implies together with elliptic regularity theory that the operator $P \circ \mathbb{L}_{s, \rho}$ is an isomorphism from $P \mathcal{C}_{s, \rho}^{2, \alpha}$ into $P \mathcal{C}^{0, \alpha}\left(\overline{B_{1}^{n}}\right)$. Here letting

$$
\left(\Phi(s, \rho, x), w^{\prime}\right):=-\int_{B_{1}^{n}} \mathcal{H}\left(s, \rho, \rho w_{1}^{s, \rho}\right) w^{\prime} d x+\oint_{\partial B_{1}^{n}} \mathcal{B}\left(s, \rho, \rho w_{1}^{s, \rho}\right) w^{\prime} d s
$$

one has

$$
\|\Phi(s, \rho, \cdot)\|_{\mathcal{C}^{0, \alpha}} \leq c \rho^{2} \quad \forall s \in[a, b] .
$$

Consequently for $\rho$ small, our fixed point problem

$$
w=\left(\left.P \circ \mathbb{L}_{\rho, s}\right|_{\mathcal{C}_{s, \rho}^{2, \alpha}}\right)^{-1}\left\{P \circ \Phi(s, \rho, x)+\rho P \circ \mathcal{Q}_{s, \rho}(w)\right\}
$$

admits a unique solution $w^{s, \rho} \in P \mathcal{C}_{s, \rho}^{2, \alpha}$, in a ball of radius $c \rho^{2}$ of $P \mathcal{C}_{s, \rho}^{2, \alpha}$. More precisely

$$
\begin{equation*}
\int_{B_{1}^{n}} w^{s, \rho} d x=0 \quad \text { and } \quad\left\|w^{s, \rho}\right\|_{\mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c \rho^{2} \quad \forall s \in[a, b] . \tag{29}
\end{equation*}
$$

Furthermore it follows from the construction that, for all $k \geq 0$

$$
\left\|\frac{\partial^{k} w^{s, \rho}}{\partial s^{k}}\right\|_{\mathcal{C}^{2, \alpha}\left(\overline{B_{1}^{n}}\right)} \leq c_{k} \rho^{2} \quad \forall s \in[a, b]
$$

for some constant $c_{k}$ which does not depend on $s$ nor on $\rho$ small. We then conclude that $P \circ \mathcal{H}\left(s, \rho, \rho w_{1}^{s, \rho}+w^{s, \rho}\right)=0$ hence the existence of a real number $\lambda_{s, \rho} \sim \rho^{2}$ such that (28) is satisfied.

We have to mention that by (29), provided $\rho$ is small, the corresponding disc $\mathscr{D}_{s, \rho}:=\mathscr{D}_{s, \rho}\left(\tilde{w}^{s, \rho}\right)$ with $\tilde{w}^{s, \rho}=\rho w_{1}^{s, \rho}+w^{s, \rho}$ is embedded into $\Omega_{\rho}$. This defines a one dimensional manifold of sets satisfying (28):

$$
\mathcal{Z}_{\rho}:=\left\{\mathscr{D}_{s, \rho} \subset \Omega_{\rho}, \partial \mathscr{D}_{s, \rho} \subset \partial \Omega_{\rho} \quad: \quad s \in[a, b]\right\}
$$

Remark 5.1 We notice that, in section 5.1, the same argument as above implies that whenever $\phi^{\prime \prime}\left(s_{0}\right)=0$, there will be a capillary disc centered at $\gamma\left(\rho s_{0}\right)$ with constant and small mean curvature.

## Variational argument:

We will show that in fact problem (26) can be reduced to a finite dimensional one. We now define the reduced functional $\varphi_{\rho}:[a, b] \rightarrow \mathbb{R}$ by

$$
\varphi_{\rho}(s):=\operatorname{Area}\left(\mathscr{D}_{s, \rho}\right)
$$

for any $\mathscr{D}_{s, \rho} \in \mathcal{Z}_{\rho}$. We have to show the following
Lemma 5.2 There exists $\rho_{0}$ small such that for any $\rho \in\left(0, \rho_{0}\right)$ if $s$ is a critical point of $\varphi_{\rho}$ then $\lambda_{s, \rho}=0$.
Proof. Let $\lambda \in \mathbb{R}$ and let $q=\gamma(\rho(s+\lambda t))$. Then provided $t$ is small, it is clear that the hyper-surface $\mathscr{D}_{q, \rho}$ can be written as a normal graph over $\mathscr{D}_{p, \rho}, p=\gamma(\rho s)$ by a smooth function $g_{p, \rho, t, \lambda}$. This defines the variation vector field

$$
\zeta_{p, \rho, \lambda}=\left.\frac{\partial g_{p, \rho, t, \lambda}}{\partial t}\right|_{t=0} N_{\mathscr{D}_{p, \rho}} .
$$

Letting $Z$ be the parallel transport of $\lambda E_{0}$ along geodesics issued from $p=\gamma(\rho s)$. Then, we can easily get the estimates:

$$
\|\zeta-Z\| \leq c \rho|\lambda| .
$$

Assume that $s$ is a critical point of $\varphi_{\rho}$ then from the first variation of area see [17],

$$
\begin{aligned}
0 & =\left.\frac{d \varphi_{\rho}(\rho(s+\lambda t))}{d t}\right|_{t=0}=\lambda \rho \varphi_{\rho}^{\prime}(s) \\
& =n \int_{\mathscr{D}_{s, \rho}} H_{\mathscr{D}_{s, \rho}}\left\langle\zeta, N_{\mathscr{D}_{s, \rho}}\right\rangle d s+\oint_{\partial \mathscr{D}_{s, \rho}}\left\langle\zeta, N_{\partial \mathscr{D}_{s, \rho}}^{\mathscr{\mathscr { s }}, \rho}\right\rangle,
\end{aligned}
$$

where $N_{\partial \mathscr{D}_{s, \rho}}^{\mathscr{D}_{s, \rho}} \in T \mathscr{D}_{s, \rho}$ stands for the normal of $\partial \mathscr{D}_{s, \rho}$ in $\mathscr{D}_{s, \rho}$. Therefore by construction one has

$$
\begin{equation*}
0=\lambda_{s, \rho} \int_{\mathscr{D}_{s, \rho}}\left\langle\zeta, N_{\mathscr{D}_{s, \rho}}\right\rangle d s \tag{30}
\end{equation*}
$$

Notice that

$$
\left\langle\zeta, N_{\mathscr{D}_{s, \rho}}\right\rangle-\left\langle Z, Y_{0}\right\rangle=\left\langle\zeta-Z, N_{\mathscr{D}_{s, \rho}}\right\rangle+\left\langle Z, N_{\mathscr{D}_{s, \rho}}-Y_{0}\right\rangle,
$$

so using the fact that $N_{\mathscr{D}_{s, \rho}}=Y_{0}+\mathcal{O}(\rho)$, see $\S$ 4.1.1, we have

$$
\left|\left\langle\zeta, N_{\mathscr{D}_{s, \rho}}\right\rangle-\lambda\right| \leq c \rho|\lambda| .
$$

Inserting this in (30), we get

$$
-\lambda \lambda_{s, \rho} \operatorname{Area}\left(\mathscr{D}_{s, \rho}\right) \leq c \rho\left|\lambda_{s, \rho}\right||\lambda| \operatorname{Area}\left(\mathscr{D}_{s, \rho}\right)
$$

but since $\operatorname{Area}\left(\mathscr{D}_{s, \rho}\right)=\operatorname{Area}\left(\rho B_{1}^{n}\right)+O_{s}\left(\rho^{2+n}\right)$ by (11); (27) and (29), it follows that

$$
-\lambda \lambda_{s, \rho} \leq c \rho\left|\lambda_{s, \rho}\right||\lambda| .
$$

Therefore taking $\lambda=-\lambda_{s, \rho}$, we see that $\left|\lambda_{s, \rho}\right|^{2} \leq c \rho\left|\lambda_{s, \rho}\right|^{2}$ and this implies that $\lambda_{s, \rho=0}$.

We shall end the proof of the Theorem 1.3 by giving the expansion of $\varphi_{\rho}$. From (29) the first fundamental form $h_{i j}$ of a disc $\mathscr{D}_{s, \rho}$ expands as

$$
\rho^{-2} h_{i j}=\delta_{i j}+\frac{\rho^{2}}{3}\left\langle R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\frac{\rho^{3}}{6}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle+\mathcal{O}_{s}\left(\rho^{4}\right),
$$

where $p=\gamma(\rho s) \in \Gamma$.
From the formula

$$
\sqrt{\operatorname{det}(I+A)}=1+\frac{1}{2} \operatorname{tr}(A)+O\left(|A|^{2}\right),
$$

we obtain the volume form:

$$
\rho^{-n} \sqrt{\operatorname{det}(h)}=1-\frac{\rho^{2}}{6}\left\langle R_{p}\left(X, E_{i}\right) X, E_{i}\right\rangle+\frac{\rho^{3}}{12}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{i}\right\rangle+\mathcal{O}_{s}\left(\rho^{4}\right)
$$

and since by oddness $\int_{B_{1}^{n}}\left\langle\nabla_{X} R_{p}\left(X, E_{i}\right) X, E_{j}\right\rangle d x=0$ we deduce that

$$
\varphi_{\rho}(s)=\operatorname{Area}\left(\mathscr{D}_{s, \rho}\right)=\operatorname{Area}\left(B_{\rho}^{n}\right)\left(1-\frac{\rho^{2}}{6 n} \sum_{i, j=1}^{n}\left\langle R_{p}\left(E_{j}, E_{i}\right) E_{j}, E_{i}\right\rangle+O_{s}\left(\rho^{4}\right)\right) .
$$

Thus setting

$$
\psi_{\rho}(s):=\frac{6 n}{\rho^{2}}\left(1-\frac{\varphi_{\rho}(s)}{\operatorname{Area}\left(B_{\rho}^{n}\right)}\right)=\sum_{i, j=1}^{n}\left\langle R_{p}\left(E_{j}, E_{i}\right) E_{j}, E_{i}\right\rangle+O_{s}\left(\rho^{4}\right)
$$

we get the result.

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